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THEORETICAL STUDY OF NON-STANDARD IMAGING CONCEPTS. VCLUME I

David L. Fried

Optical Science Consultants

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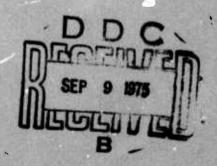
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instant will be almost diffraction-limited. (The problem is perhaps most succinctly defined by the question — How many pictures do you have to take to get a good ore?) The numerical results show that the probability is an exponential function of aperture area divided by r^2 . If D/r = (7, 10, 15), the probability of getting a nearly diffraction-limited image is found to be $P_{\text{CENSOR}} \approx (3 \times 10^{-3} \ , \ 1 \times 10^{-6} \ , \ 3.4 \times 10^{-16})$. The functional relationship is

 $P_{CENSOR} \approx 5.6 \exp [-0.1557 (D/r_0)^2]$

Derivation and basic results are presented in this main volume. Certain of the more voluminous tables are presented in a separate addendum volume.

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THEORETICAL STUDY OF NON-STANDARD IMAGING CONCEPTS

Dr. David L. Fried

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ABSTRACT

In Part I of this report, the performance of a CENSORING system is examined from the point of view of determining the probability, Pcensor, that at any instant of time the random wavefront distortion over a circular aperture of diameter D will be small enough to allow a nearly diffractionlimited image to be formed. It is pointed out that the effective wavefront distortion for this calculation is not the total distortion function, $\phi(\vec{r})$, but rather $\varphi(\vec{r};D)$, which represents the distortion after subtraction of the instantaneous average phase and tilt over the aperture. The problem of calculating the probability is related to the probabilities for the value of the various random coefficients β_n of the decomposition of $\phi(\vec{r};D)$ into a series β_n $f_n(\vec{r};D)$, where the functions f_n are orthonormal functions over the aperture chosen so as to make the various β_n statistically independent. (It is noted that the β_n are gaussianly distributed, since ϕ is a gaussian random function.) The appropriate Karhunen-Loeve homogeneous integral equation is developed to allow the f, to be obtained as eigenfunctions. The variance of the B, are seen to be the corresponding eigenvalues. It is then shown how the probability Pcensor can be calculated as a multi-dimensional integral over the product of gaussian distribution with the variances corresponding to those of the β_n . The numerical solution of the Karhunen-Loeve homogeneous integral equation to obtain the eigenvalues (and the eigenfunctions) and the numerical evaluation of the multi-dimensional integral which finally gives Pcensor are taken up in Part II of this report.

In Part II of this report, the eigenvalues and the eigenfunctions for average tilt and average phase suppressed wavefront distortion on a circular aperture are developed. (The eigenvalues and eigenfunctions without tilt distortion suppressed are also developed.) Using the eigenvalue set, the CENSORING system probability, P_{CENSOR} , of getting a short exposure

image with less than one radian squared distortion averaged over the aperture is evaluated as a multi-dimensional integral, evaluated by Monte Carlo techniques. The results are found to have the form

$$P_{CENSOR} \approx 5.6 \exp [-0.1557 (D/r_0)^2]$$

The exponential dependence on aperture area is in agreement with an earlier conjecture by Hufnagel (though with significantly different coefficients than were suggested by Hufnagel).

It is noted that this exponential dependence on aperture diameter makes the proper selection of D/r_0 a very critical aspect of a CENSOR-ING experiment. It is pointed out that if independent samples of wavefront distortion are obtained every 10 msec, and $D/r_0 = 15$ (corresponding to $D = 1.5 \, \mathrm{m}$, $r_0 = 0.1 \, \mathrm{m}$), then we would have to wait about 800 million hours for a good picture. If D/r_0 is reduced to a value of 10 (corresponding to $D = 1.0 \, \mathrm{m}$, $r_0 = 0.1 \, \mathrm{m}$), the waiting time is reduced to about 2.5 hours, while if D/r_0 is reduced to 7 (corresponding to $D = 0.7 \, \mathrm{m}$, $r_0 = 0.1 \, \mathrm{m}$), the waiting time drops to only 3.5 seconds.

The report has been bound in two volumes. The main volume presents the derivation and principal results. The addendum volume presents the more voluminous tables of intermediate results, particularly those that may be of use in working other problems.

PART I

Formal Theory

o f

CENSORING Systems Operation

Introduction

The concept of CENSORING as a method of obtaining high resolution images through atmospheric turbulence is based on the assumption that of all possible forms that random wavefront distortion will assume during some period of time, there is a finite probability that at some instant the random pattern will very nearly resemble a plane wave. At that instant, a nearly diffraction-limited image can be obtained. A CENSORING system would be able to recognize this condition quickly enough to allow a photograph to be taken just then, while preventing exposures at times of more normal distortion.

The simplicity of the CENSORING concept makes it seem quite attractive, but on the other hand it is entirely dependent on random occurrences for its operation. It is therefore critical that we understand the probabilities involved and be able to estimate the time we will have to wait, on ar average, before the CENSORING system will provide us with a picture. It appears likely that the probability that at any instant of time the distorted wavefront will be reasonably close to a plane wave is inversely proportional to the telescope aperture area (divided by ${\bf r_0}^2$) so that there is a practical limit of useful telescope size for a CENSORING system. The larger the telescope diameter, the longer we have to wait for a good picture, with waiting time increasing dramatically as telescope size is increased.

For this reason, we desire a quantitative understanding of the probabilities involved in a CENSORING system's operation. This paper is aimed at the formulation of that theory. Here we shall be concerned with setting up the basic formulations and deriving equations suitable for computer evaluation. In a later paper, we shall carry out the necessary computer calculations.

Wavefront Distortion Analysis

The key to analysis of the wavefront distortion probabilities involved in CENSO ING operation is the decomposition of a sample of the random wavefront taken over the aperture into a set of orthonormal functions whose coefficients in the decomposition series representation are independent random variables. From knowledge of the mean-square value of these random coefficients, we can calculate the appropriate probabilities for CENSORING system operation.

The orthonormal decomposition with independent random coefficients is related to the Karhunen-Loeve theorem and we can anticipate that the development of the orthonormal functions and the evaluation of the mean-square value of the coefficients will depend on the solution of a homogeneous integral equation for its eigenfunctions and eigenvalues, respectively. The key to that effort is the development of the kernel for the integral equation. The kernel is developed from the statistics of wavefront distortion and, as we may naturally expect, is related to the phase structure function. However, as we shall see, the relationship is by no means trivial and requires careful development.

In order to calculate the kernel, we first have to define in exact terms the nature of the wavefront distortion statistics and the "portion" of the distortion that is of concern to us. We denote the random wavefront distortion (measured in radians of phase) at a point \vec{r} on the aperture plane by $\phi(\vec{r})$. Over a circular aperture of diameter D a random sample of the distorted wavefront has a random average phase $\vec{\phi}$, and a random average tilt $\vec{\alpha}$, where

$$\vec{\phi} = (\frac{1}{4} \pi D^2)^{-1} \int d\vec{r} W(\vec{r}, D) \phi(\vec{r}) , \qquad (1)$$

and

$$\vec{\alpha} = (\frac{1}{64} \pi D^4)^{-1} \int \vec{dr} W(\vec{r}, D) \vec{r} \phi(\vec{r}) , \qquad (2)$$

where $W(\vec{r}, D)$ is an aperture function defining, in the \vec{r} -plane, a circle of diameter D centered at the origin. This aperture function is defined by the equation

$$W(\vec{r}, D) = \begin{cases} 1 & \text{if } |\vec{r}| \leq \frac{1}{2}D \\ 0 & \text{if } |\vec{r}| > \frac{1}{2}D \end{cases} , \qquad (3)$$

so that in effect, it defines the limits of integration for the \vec{r} -integration in Eq.'s (1) and (2), which are otherwise taken to be over the infinite \vec{r} -plane. The normalization in Eq. (2) has been chosen so that if $\phi(\vec{r}) \equiv \vec{a} \cdot \vec{r}$, then we would obtain from Eq. (2) the relationship $\vec{\alpha} = \vec{a}$.

We note that in forming a short-exposure image, neither the average phase, $\overline{\phi}$, nor the wavefront tilt, $\overline{\alpha}$, disturb the resolution of the image. The tilt produces an image shift which is basically a nonobservable in the sense that without special effort to provide an absolute angular orientation reference, the effect of the tilt will not be measurable. Thus, from our point of view, the effective instantaneous random wavefront distortion over the aperture is

$$\varphi(\vec{r};D) = \varphi(\vec{r}) - \overline{\varphi} - \vec{\alpha} \cdot \vec{r} \qquad . \tag{4}$$

If φ is small enough, then the image will be nearly diffraction-limited no matter how large $\overline{\varphi}$ and $\overrightarrow{\alpha}$ are. We wish to calculate the probability distribution for φ as the basis for determining the statistics of the operation of a CENSORING system.

To provide the basis for our calculations of these statistics, at this point we introduce the set of functions, $\{f_n(\vec{r};D)\}$ with the orthonormal property that

$$\int d\vec{r} W(\vec{r}, D) f_n^*(\vec{r}; D) f_n(\vec{r}; D) = \begin{cases} 1 & \text{if } n = n' \\ 0 & \text{if } n \neq n' \end{cases}$$
 (5)

and the completeness property that for any random sample $\phi(\vec{r};D)$ we can write

$$\varphi(\vec{r};D) = \sum_{n} \beta_{n} f_{n}(\vec{r};D) \qquad (6)$$

Where $\{\beta_n\}$ is an appropriately chosen set of coefficients. Because of the orthonormal property of f_n , as defined by Eq. (5), it follows from Eq. (6) that

$$\beta_{\mathbf{n}} = \int d\vec{\mathbf{r}} \ W(\vec{\mathbf{r}}, D) \ f_{\mathbf{n}}^{*}(\vec{\mathbf{r}}; D) \ \varphi(\vec{\mathbf{r}}; D) \ . \tag{7}$$

Obviously, then, just as $\,^{\phi}$ is a random function, $\,^{\beta}_n$ is a random variable. We also note that since $\,^{\phi}$ is a gaussian random function, then $\,^{\phi}$ and $\,^{\gamma}$, being linear functions of $\,^{\phi}$ are gaussian random variables. From this, in turn, it follows that $\,^{\phi}$ is a gaussian random function -- and this, in turn, implies that $\,^{\beta}_n$, being a linear function of $\,^{\phi}$, is a gaussian random variable. This fact, together with our ability to calculate the variance of $\,^{\beta}_n$ (which we shall obtain as the eigenvalues of the integral equation defining $\,^{\epsilon}_n$), will provide the basis of calculating the probability of $\,^{\phi}$ taking a low enough wavefront distortion form.

The key to the definition of the set of orthonormal functions $\{f_n\}$ from all possible sets of functions that are orthonormal over the region defined by $W(\vec{r}, D)$ is the requirement that the various random coefficients β_n must be independent. We require that

$$\langle \beta_n^* \beta_{n'} \rangle = \begin{cases} B_n^2(D) & \text{if } n = n' \\ 0 & \text{, if } n \neq n' \end{cases} , \tag{8}$$

where $B_n^{\ 2}(D)$ denotes the variance of the random variable ρ_n . (In writing $B_n^{\ 2}$, we have chosen, as a matter of convenience for later work, to make the dependence of $B_n^{\ 2}$ on the aperture diameter, D, explicit.) To see the

implications of Eq. (8), we consider the quantity

$$\delta = \langle \int d\vec{r}' W(\vec{r}', D) \varphi^*(\vec{r}'; D) \varphi(\vec{r}, D) f_{\mathbf{n}}(\vec{r}'; D) \rangle . \qquad (9)$$

We define the covariance of ϕ as

$$C_{\varphi}(|\vec{r} - \vec{r}'|;D) = \langle \varphi^*(\vec{r};D) \varphi(\vec{r}';D) \rangle , \qquad (10)$$

where the homogeneity and isotropy of the propagation statistics have allowed us to write the dependence on \vec{r} and \vec{r}' in the form of $|\vec{r} - \vec{r}'|$. We note that by interchanging the order of ensemble averaging and integration, we can rewrite Eq. (9) in the form

$$\mathbf{s} = \int d\vec{\mathbf{r}}' \ W(\vec{\mathbf{r}}', D) \ C_{\varphi} \ (|\vec{\mathbf{r}} - \vec{\mathbf{r}}'|; D) \ f_{n}(\vec{\mathbf{r}}'; D) \quad . \tag{11}$$

However, if we use Eq. (6) to provide a series representation for $\phi^*(\vec{r}';D)$ to be substituted into Eq. (9), we get

$$\boldsymbol{\delta} = \langle \int d\vec{r}' \ W(\vec{r}', D) \ \varphi(\vec{r}; D) \ \sum_{n'} \beta_{n'} * f_{n'} * (\vec{r}'; D) \ f_{n}(\vec{r}'; D) \rangle , \qquad (12)$$

which on interchanging the order of integration and summation gives us

$$\mathcal{S} = \langle \varphi(\vec{r}; D) \sum_{n'} \beta_{n'} * \int d\vec{r}' W(\vec{r}', D) f_{n'} * (\vec{r}'; D) f_{n}(\vec{r}'; D) \rangle . \quad (13)$$

The orthonormal property of f_n as expressed in Eq. (5) allows the integration to be performed, and the result of this allows us to reduce the summation over n' to the single term for which n' = n. Thus we get

$$\delta = \langle \varphi(\vec{r}; D) \beta_n^* \rangle \qquad (14)$$

Now it we again use Eq. (6) to provide a series representation for $\phi(\vec{r};D)$, we get

$$\mathcal{E} = \langle \sum_{\mathbf{n'}} \beta_{\mathbf{n'}} f_{\mathbf{n'}}(\mathbf{r}; \mathbf{D}) \beta_{\mathbf{n}}^* \rangle$$
 (15)

An interchange of the order of summation and ensemble averaging leads to the result

$$\mathcal{S} = \sum_{\mathbf{n'}} f_{\mathbf{n'}}(\vec{\mathbf{r}}; \mathbf{D}) \langle \beta_{\mathbf{n}}^* \beta_{\mathbf{n'}} \rangle \qquad (16)$$

Now we make use of the requirement that the random coefficients β_n and β_n , be independent, as expressed by Eq. (8), which allows us to reduce the summation on n' to a single term with n' = n. Thus we get as a consequence of the requirement of independence of the β_n 's

$$\delta = B_{\bullet}^{2}(D) f_{\bullet}(\vec{r}; D) , \qquad (17)$$

which, when combined with Eq. (11), gives the basic Karhunen-Loeve homogeneous integral equation

$$\int d\vec{r}' W(\vec{r}', D) C_{\varphi}(|\vec{r} - \vec{r}'|; D) f_n(\vec{r}'; D) = B_n^2(D) f_n(\vec{r}; D) . \qquad (18)$$

The eigenfunctions of this equation are the orthonormal functions we wish to work with. These functions have statistically independent coefficients in a series representative of ϕ as required by Eq. (8). The variance of these coefficients are the corresponding eigenvalues of the Karhunen-Loeve integral equation.

Our basic remaining tasks in this paper are 1) to see how the probability of the effective wavefront distortion, ϕ , being adequately small can be calculated from the eigenfunctions and eigenvalues defined by Eq. (18),

and 2) to develop an expression for C_{φ} to be substituted into Eq. (18). However, before delving too deeply into these matters, we first wish to consider the question of how we can "dedimensionalize" our problem so that a single case solution of the Karhunen-Loeve integral equation will provide the basic treatment for all aperture diameters and strength of turbulence conditions.

Dedimensionalization

The basic statistics of wavefront distortion are provided by the wave-structure function, * $\mathfrak{F}(r)$, where

$$\mathcal{L}(|\vec{r} - \vec{r}'| = \langle |\phi(\vec{r}) - \phi(\vec{r}')|^2 \rangle \qquad (19)$$

It can be shown that the value of the wave-structure function may be written as

$$\mathcal{S}(r) = 6.88 (r/r_0)^{5/3} , \qquad (20)$$

where r_0 is a quantity with the dimensions of length. The value of r_0 is determined by the optical wavelength in question and the distribution of the strength of turbulence along the propagation path. For our purposes here, the nature of that relationship is of no consequence. It is sufficient to know the value of r_0 as this quantity, as we shall see, completely characterizes the effective strength of turbulence for evaluation of the performance of a CENSORING system.

At this point, we state without proof the fact that we can extract the dependence of $C_{\phi}(|\vec{r}-\vec{r}'|;D)$ on both the strength of turbulence, as

^{*} It should be noted that although we have spoken of ϕ as a phase distortion, we actually consider it to be a complex phase with its real part corresponding to ordinary phase and the negative of its imaginary part corresponding to log-amplitude variations. Thus both ϕ and the random coefficients $\{\beta_n\}$ are complex quantities. In all of our analysis, we have been careful to introduce complex conjugation where appropriate, though we have not made a point of the fact that the quantities are complex. In most cases, the imaginary part is much smaller than the real part.

defined by r_0 , and the aperture diameter, D , by writing

$$C_{\phi}(|\vec{r} - \vec{r}'|; D) = (D/r_0)^{5/3} \mathfrak{C}(|\vec{x} - \vec{x}'|)$$
, (21)

where

$$\vec{x} = \vec{r}/D$$
 and $\vec{x}' = \vec{r}'/D$. (22)

We note in particular that $\mathfrak{F}(|\vec{x}-\vec{x}'|)$ is independent of the value of D. The validity of Eq. (21) will be established in a later section, where we shall consider in detail the evaluation of C_{ϕ} and will give an explicit expression for \mathfrak{F} .

If we substitute Eq. (21) into Eq. (18) and change the variables from \vec{r} , \vec{r} to \vec{x} , \vec{x} , noting that $d\vec{r}$ goes into $D^2 d\vec{x}$, we get the result that

$$\int d\vec{x}' W(\vec{x}', 1) \mathfrak{T}(|\vec{x} - \vec{x}'|) \mathfrak{F}_n(\vec{x}'; D) = \mathfrak{R}_n^2(D) \mathfrak{F}_n(\vec{x}; D) , \qquad (23)$$

where

$$\mathfrak{F}_{\mathbf{n}}(\mathbf{x}; \mathbf{D}) \equiv \mathbf{f}_{\mathbf{n}}(\mathbf{D}\mathbf{x}; \mathbf{D}) , \qquad (24)$$

and

$$\mathfrak{A}_{n}^{S}(D) = D^{-2} (D/r_{0})^{-6/3} B_{n}^{S}(D)$$

$$= D^{-11/3} r_{0}^{6/3} B_{n}^{S}(D)$$
(25)

We note that Eq. (23) is a homogeneous integral equation with eigenfunction $\mathfrak{F}_n(\vec{x};D)$ and eigenvalue $\mathfrak{B}_n^2(D)$. However, since the kernel of that integral equation, as well as the limits on the integration, is independent of the aperture diameter, D, then it follows that the eigenfunctions and eigenvalues must be independent of D. We therefore may write the eigenfunction $\mathfrak{F}_n(\vec{x};D)$ as $\mathfrak{F}_n(\vec{x})$ without any loss of definiteness, with the understanding

that the eigenfunction we originally sought, i.e., $f_n(\vec{r};D)$ can be written as

$$f_n(\vec{r};D) = g_n(\vec{r}/D) \qquad . \tag{26}$$

Since the eigenvalues of Eq. (23), i.e., $\{\mathfrak{B}_n^{\;2}(D)\}$, are independent of D (and of r_0), we can without any loss of information write them as $\{\mathfrak{F}_n^{\;2}\}$. It then follows that the eigenvalues we originally sought, namely, $B_n^{\;2}(D)$, can be written as

$$B_n^2(D) = D^{11/5} r_0^{-5/5} \mathfrak{R}_n^2$$
 (27)

We recall that in accordance with the above discussion, $\{\mathfrak{B}_{n}^{2}\}$ and $\{\mathfrak{F}_{n}(\vec{x})\}$ are the set of eigenvalues and eigenfunctions for the Karhunen-Loeve homogeneous integral equation

$$\int d\vec{x}' W(\vec{x}', 1) \, \mathfrak{G}([\vec{x} - \vec{x}']) \, \mathfrak{F}_n(\vec{x}') = \mathfrak{B}_n^2 \, \mathfrak{F}_n(\vec{x}) \quad . \tag{28}$$

At this point, we need only develop an expression for $\mathfrak{F}(|\vec{x}-\vec{x}'|)$, in accordance with Eq. (21), to set up the problem for solution of the integral equation in its general form. Then, after obtaining these generalized solutions, for any value of aperture diameter, D, and turbulence parameter, r_0 , we can obtain the eigenvalues and eigenfunctions, $\{B_n^{(2)}(D)\}$ and $\{f_n(\vec{r};D)\}$ from Eq.'s (26) and (27), for that particular problem. In the next section, we take up the evaluation of \mathfrak{F} .

Evaluation of the Kernel

Though we are interested in developing the kernel, $\mathfrak{T}(|\vec{x}-\vec{x}'|)$, we shall proceed in that by means of Eq. (21), and in most of this section shall be concerned with the evaluation of the original kernel, $C_{\phi}(|\vec{r}-\vec{r}'|;D)$, as defined by Eq. (10). We shall find, at the end of this section, that extraction of a result for $\mathfrak{T}(|\vec{x}-\vec{x}'|)$ then drops out as a trivial additional manipulation.

If we substitute Eq. (4) into Eq. (10), we obtain

$$C_{\varphi}(|\vec{r} - \vec{r}'|; D) = \langle [\varphi(\vec{r}) - \overline{\varphi} - \vec{\alpha} \cdot \vec{r}]^* [\varphi(\vec{r}') - \overline{\varphi} - \vec{\alpha} \cdot \vec{r}'] \rangle$$

$$= \langle \phi^*(\vec{r}) \varphi(\vec{r}') \rangle - \langle \phi^*(\vec{r}) \overline{\varphi} \rangle - \langle \varphi(\vec{r}') \overline{\varphi}^* \rangle + \langle \overline{\varphi}^* \overline{\varphi} \rangle$$

$$- \langle \phi^*(\vec{r}) \vec{\alpha} \cdot \vec{r}' \rangle - \langle \varphi(\vec{r}') \vec{\alpha}^* \cdot \vec{r} \rangle + \langle \overline{\varphi}^* \vec{\alpha} \cdot \vec{r}' \rangle + \langle \overline{\varphi}^* \vec{\alpha} \cdot \vec{r}' \rangle$$

$$+ \langle \vec{\alpha}^* \cdot \vec{r} \vec{\alpha} \cdot \vec{r}' \rangle \qquad (29)$$

Making use of Eq.'s (1) and (2), and interchanging the order of ensemble averaging and integration, we can rewrite Eq. (29) in the form

$$C_{\phi}(|\vec{r} - \vec{r}'|; D) = \langle \phi^*(\vec{r}) \phi(\vec{r}') \rangle$$

$$- (\frac{1}{2} \pi D^2)^{-1} \int d\vec{r}'' W(\vec{r}', D) \langle \phi^*(\vec{r}) \phi(\vec{r}'') \rangle$$

$$- (\frac{1}{4} \pi D^2)^{-1} \int d\vec{r}'' W(\vec{r}', D) \langle \phi^*(\vec{r}') \phi(\vec{r}'') \rangle$$

$$+ (\frac{1}{4} \pi D^2)^{-2} \int d\vec{r}'' d\vec{r}'' W(\vec{r}'', D) W(\vec{r}'', D)$$

$$\times \langle \phi^*(\vec{r}'') \phi(\vec{r}''') \rangle$$

$$- (\frac{1}{64} \pi D^4)^{-1} \int d\vec{r}'' W(\vec{r}'', D) \langle \phi^*(\vec{r}) \phi(\vec{r}'') \rangle \vec{r}'' \cdot \vec{r}'$$

$$- (\frac{1}{64} \pi D^4)^{-1} \int d\vec{r}'' W(\vec{r}'', D) \langle \phi^*(\vec{r}'') \phi(\vec{r}') \rangle \vec{r}'' \cdot \vec{r}''$$

$$+ (\frac{1}{16} \pi D^2)^{-2} \int d\vec{r}'' d\vec{r}'' W(\vec{r}'', D) W(\vec{r}''', D)$$

$$\times \langle \phi^*(\vec{r}'') \phi(\vec{r}''') \rangle \vec{r}' \cdot \vec{r}'''$$

$$+ (\frac{1}{64} \pi D^4)^{-2} \int d\vec{r}'' d\vec{r}''' W(\vec{r}'', D) W(\vec{r}''', D)$$

$$\times \langle \phi^*(\vec{r}''') \phi(\vec{r}'''') \rangle \vec{r}' \cdot \vec{r}'''$$

$$+ (\frac{1}{64} \pi D^4)^{-2} \int d\vec{r}'' d\vec{r}''' W(\vec{r}'', D) W(\vec{r}''', D)$$

$$\times \langle \phi^*(\vec{r}''') \phi(\vec{r}'''') \rangle (\vec{r} \cdot \vec{r}''') (\vec{r}' \cdot \vec{r}'''') . \qquad (30)$$

We note that in the first four terms on the right-hand-side of Eq. (30), if the ensemble averages are each replaced by a constant, then the sum of

the four terms vanishes. This means that we may add or subtract a constant value from the ensemble averages in these terms. We further note that if the ensemble average in any of the remaining terms is replaced by a constant, the integration is such that the term will vanish. Hence we may add or subtract a constant from any of these ensemble averages. Because of the stationarity of $\langle \phi^*(\vec{r}) \phi(\vec{r}) \rangle$, it is identical in value to $\langle \phi^*(\vec{r}') \phi(\vec{r}') \rangle$ or $\langle \phi^*(\vec{r}'') \phi(\vec{r}'') \rangle$, etc. We find then that by appropriate manipulation (which, amongst other things, includes taking note of the fact that for physical reasons C_{ϕ} is real, so that we can replace the right-hand-side of Eq. (30) with one-half the sum of its value plus its complex conjugate), we can replace each ensemble average of the product of phases on the right-hand-side of Eq. (30) by minus one-half the ensemble average of the difference of the phases absolute value square. Thus

$$\langle \phi^*(\vec{r}) \phi(\vec{r}') \rangle \Rightarrow -\frac{1}{2} \mathcal{B}(|\vec{r} - \vec{r}'|)$$
 (31)

or

$$\langle \phi^*(\vec{r}'') \phi(\vec{r}''') \rangle \Rightarrow -\frac{1}{2} \mathcal{D}(|\vec{r}'' - \vec{r}'''|) , \qquad (32)$$

etc., where \mathcal{L} , the wave-structure function, is defined in Eq. (19), with the value given in Eq. (20).

We can now rewrite Eq. (30) in the form

$$C_{\varphi}(|\vec{r}-\vec{r}'|;D) = T_1 + T_2 + T_3 + T_4 + T_5 + T_6 + T_9$$
, (33)

where

$$T_1 = -\frac{1}{2} \mathcal{B} \left(\left| \vec{r} - \vec{r}' \right| \right) \qquad , \tag{34}$$

$$T_{2} = \frac{1}{2} \left(\frac{1}{4} \pi D^{2} \right)^{-1} \int d\vec{r} W(\vec{r}, D) \mathcal{B}(|\vec{r} - \vec{r}|) , \qquad (35)$$

$$T_{3} = \frac{1}{2} \left(\frac{1}{4} \pi D^{2} \right)^{-1} \int d\vec{r} W(\vec{r}, D) \mathcal{D}(|\vec{r}' - \vec{r}''|) , \qquad (36)$$

$$T_{4} = -\frac{1}{2} \left(\frac{1}{4} \pi D^{2} \right)^{-2} \iint d\vec{r} \, \vec{w} \, (\vec{r} \, \vec{r}, D) \, W(\vec{r} \, \vec{m}, D)$$

$$\times \, \mathcal{D}(|\vec{r} \, \vec{r} \, - \vec{r} \, \vec{m}|) \qquad , \qquad (37)$$

$$T_{5} = \frac{1}{2} \left(\frac{1}{64} \pi D^{4} \right)^{-1} \int d\vec{r} W(\vec{r}, D) \mathcal{B}(|\vec{r} - \vec{r}|) \vec{r} \vec{r} \vec{r}, \quad (38)$$

$$T_{a} = \frac{1}{2} \left(\frac{1}{64} \pi D^{4} \right)^{-1} \int d\vec{r} W(\vec{r}, D) \mathcal{L}(|\vec{r}' - r''|) \vec{r} \vec{r} , \qquad (39)$$

$$T_{g} = -\frac{1}{2} \left(\frac{1}{64} \pi D^{4} \right)^{-2} \iint d\vec{r} \, d\vec{r} \, W(\vec{r}, D) \, W(\vec{r}, D)$$

$$\times \, \mathcal{D}(|\vec{r} - \vec{r}|) \, (\vec{r} \cdot \vec{r}) \, (\vec{r} \cdot \vec{r}) \, . \tag{40}$$

In developing Eq. (33) from Eq. (29), we have dropped the two terms

$$T_{7} = -\frac{1}{2} \left(\frac{1}{16} \pi D^{3} \right)^{-2} \iint d\vec{r} \, \vec{m} \, \vec{dr} \, \vec{m} \, W(\vec{r} \, \vec{m}, D) \, W(\vec{r} \, \vec{m}, D)$$

$$\times \mathcal{L}(|\vec{r} \, \vec{m} - \vec{r} \, \vec{m}|) \, \vec{r}' \cdot \vec{r} \, \vec{m} \qquad (41)$$

$$T_{8} = -\frac{1}{2} \left(\frac{1}{16} \pi D^{3} \right)^{-2} \iint d\vec{r} \, d\vec{r} \, W(\vec{r} \, D) \, W(\vec{r} \, D)$$

$$\times \mathcal{P}(\vec{r} \, -\vec{r} \, D) \, \vec{r} \cdot \vec{r} \, D \qquad (42)$$

which correspond to the seventh and eighth terms in Eq. (29). Our reason for doing this is that both terms can be shown to have zero value. To see that these terms do indeed vanish, we note that if we consider pairs of values of \vec{r} and \vec{r} that hold \vec{r} , \vec{r} , and the angle between \vec{r} and r constant, then as we integrate about the orientation of the r -vector, everything in the integrand is constant except $\vec{r} \cdot \vec{r}$. As a result of the variation of this factor, the integration over 2m yields a zero value. (It is interesting to note that T_{γ} and T_{θ} correspond to the correlation of $\overline{\phi}$ and $\overset{
ightarrow}{\sigma}$, which, as we should have expected, are uncorrelated.)

With Eq. (33) established, we now look into the problem of using Eq. (20) for β to simplify our expression for C_{ϕ} . It is convenient at this point to introduce the following functions:

$$\mathfrak{G}_{0}(x) = 3.44 x^{5/3} , \qquad (43)$$

$$\mathfrak{G}_{1}(x) = 3.44 \left(\frac{1}{4}\pi\right)^{-1} \int_{0}^{1/2} dx x x \int_{0}^{2\pi} d\theta x (x^{2} + x^{2} - 2xx^{2}\cos\theta)^{5/6} \qquad (44)$$

$$\mathfrak{G}_{2} = 8 \int_{0}^{1/2} dx x x \mathfrak{G}_{1}(x^{2}) , \qquad (45)$$

(45)

$$G_{3}(x) = 3.44 \left(\frac{1}{64}\pi\right)^{-1} \int_{0}^{1/2} dx \, x^{2} \int_{0}^{2\pi} d\theta \, \cos \theta \, d\theta$$

$$\times \left(x^{2} + x^{2} - 2xx \, \cos \theta^{2}\right)^{5/6} \qquad (46)$$

$$\mathfrak{G}_{4} = 64 \int_{0}^{1/2} dx'' x'''^{2} \mathfrak{G}_{3}(x'') . \tag{47}$$

We shall see shortly that C_{φ} can be expressed in terms of these five functions (or more precisely, three functions and two constants).

From Eq. (20), we see that

$$T_1 = -3.44 (r^2 + r'^2 - 2r r' \cos \theta')^{5/6} r_0^{-5/3}$$

$$= -3.44 (D/r_0)^{5/3} [(r/D)^2 + (r'/D)^2 - 2(r/D)(r'/D) \cos \theta']^{5/6}$$

$$= -(D/r_0)^{5/3} \mathcal{O}_0 (|(\vec{r}/D) - (\vec{r}'/D)|) \qquad (48)$$

Proceeding in the same vein, we see that

$$T_{2} = 3.44 \left(\frac{1}{4}\pi\right)^{-1} \left(D/r_{0}\right)^{5/3} \int_{0}^{1/2} dx'' x'' \int_{0}^{2\pi} d\theta''$$

$$\times \left[(r/D)^{2} + x''^{2} - 2(r/D) x'' \cos \theta'' \right]^{5/6}$$

$$= (D/r_{0})^{5/3} \mathcal{O}_{1}(r/D) , \qquad (49)$$

where we have made the transformation of variables

$$\vec{r} = D \vec{x}$$
 and $d\vec{r} = D^2 d\vec{x}$, (50)

and later will also use

$$\vec{r} = D \vec{x} \quad \text{and} \quad d\vec{r} = D^2 d\vec{x} \quad . \tag{51}$$

For T_3 , working in exactly the same way but with \vec{r} replaced by \vec{r}^\prime , we get

$$T_3 = (D/r_0)^{\varepsilon/3} \otimes_1 (r'/D) \qquad . \tag{52}$$

Working in the same manner with $T_{\mathbf{4}}$, as defined in Eq. (37), we get

$$T_{4} = -3.44 \left(\frac{1}{4}\pi\right)^{-2} \left(\frac{D}{r_{0}}\right)^{5/3} \int_{0}^{1/2} dx^{*} x^{*''} \int_{0}^{1/2} d\theta^{*''} \int_{0}^{1/2} dx^{*''} x^{*''} \int_{0}^{2\pi} d\theta^{*''}$$

$$\times \left[x^{*''^{2}} + x^{*''^{2}} - 2x^{*''} x^{*''} \cos\left(\theta^{*''} - \theta^{*''}\right)\right]^{5/6}$$

$$= -3.44 \left(\frac{1}{4}\pi\right)^{-2} \left(\frac{D}{r_{0}}\right)^{5/3} \int_{0}^{1/2} dx^{*''} x^{*''} 2\pi \left\{\int_{0}^{1/2} dx^{*''} x^{*'''} \int_{0}^{2\pi} d\theta^{*''} d\theta^{*'''}\right\}$$

$$\times \left[x^{*''^{2}} + x^{*''^{2}} - 2x^{*''} x^{*''} \cos\left(\theta^{*''}\right)\right]^{5/6} \right\} \qquad (53)$$

Here we have replaced the variable θ''' by $\theta''' + \theta'''$ (treating θ''' as a constant for the θ''' -integration), and then shifted the limits of that integration from $\theta'' \leftrightarrow 2\pi + \theta'''$ to $0 \leftrightarrow 2\pi$. This then allowed the θ'' -integration to be performed, yielding a factor of 2^{π} for the final result in Eq. (53). We note that the quantity in the curly brackets in Eq. (53) is directly related to $\mathfrak{G}_1(x'')$. Thus we can write

$$T_{4} = -8 (D/r_{0})^{6/3} \int_{0}^{1/2} dx'' x'' \Theta_{1}(x'')$$

$$= -(D/r_{0})^{6/3} \Theta_{2} \qquad (54)$$

For T_5 as defined by Eq. (38), we have the necessary extra powers of D^{-1} in front of the integral to convert $\vec{r}'' \cdot \vec{r}'$ in the integrand into $\vec{x}'' \cdot \vec{x}'$. Following the same procedure as above, this allows us to

write

$$T_{5} = 3.44 \left(\frac{1}{64} \pi\right)^{-1} \left(D/r_{0}\right)^{5/3} \int_{0}^{1/2} dx'' x'' \int_{0}^{2\pi} d\theta''' x'' x' \cos \left(\theta'' - \theta'\right)$$

$$\times \left[(r/D)^{2} + x''^{2} - 2x x'' \cos \theta'' \right]^{5/6} , \qquad (55)$$

where the θ value is referenced to the \vec{r} orientation so that θ = 0 when \vec{x} is parallel to \vec{r} (or rather when \vec{r} is parallel to \vec{r}). θ is similarly referenced to the \vec{r} vector. At this point, we note that since we can rewrite $\cos(\theta^* - \theta')$ in the form

$$\cos (\theta'' - \theta') = \cos \theta'' \cos \theta' + \sin \theta'' \sin \theta' , \qquad (56)$$

then we can separate the θ'' -integration in Eq. (55) into the sum of two integrations. The first would be multiplied by a factor of $\cos\theta'$ and would have $\cos\left(\theta''-\theta'\right)$ in the integrand of Eq. (55) replaced by $\cos\theta''$. The second would be multiplied by $\sin\theta'$ and would have $\cos\left(\theta'''-\theta'\right)$ in the integrand replaced by $\sin\theta''$. We note, however, that since the integrand in the second integration would be odd in θ''' , the second integral being over the range $0\leftrightarrow2\pi$ will have zero value. Thus we obtain the result that

$$T_{5} = 3.44 \left(\frac{1}{64}\pi\right)^{-1} \left(D/r_{0}\right)^{5/3} x' \cos \theta' \int_{0}^{1/2} dx'' x''^{2} \int_{0}^{2\pi} d\theta'' \cos \theta''$$

$$\times \left[(r/D)^{2} + x''^{2} - 2x x'' \cos \theta'' \right]^{5/3}$$

$$= (D/r_{0})^{5/3} (r'/D) \cos \theta' G_{5} (r/D) . \tag{57}$$

Exactly the same procedure applied to the evaluation of T_{δ} leads to the result

$$T_6 = (D/r_0)^{6/3} (r/D) \cos \theta' \Theta_3 (r'/D)$$
 (58)

The evaluation of T_9 follows the same general approach as the above. In this case, we have enough extra powers of D^{-1} in front of the integral to allow us to convert $(\vec{r} \cdot \vec{r}'') (\vec{r}' \cdot \vec{r}''')$ into $[(\vec{r}/D) \cdot \vec{x}''][(\vec{r}'/D) \cdot \vec{x}''']$. Thus we can write

$$T_{\theta} = -3.44 \left(\frac{1}{64}\pi\right)^{-2} \int_{0}^{1/2} dx'' x'' \int_{0}^{2\pi} d\theta'' \int_{0}^{1/2} dx''' x''' \int_{0}^{2\pi} d\theta''' (r/D)(r'/D) x'' x'''$$

$$\times \cos \theta'' \cos (\theta''' - \theta'') \left[x'''^2 + x'''^2 - 2x'' x''' \cos (\theta''' - \theta'') \right]^{5/6}$$

$$= -3.44 \left(\frac{1}{64}\pi\right)^{-2} (r/D)(r'/D) \int_{0}^{1/2} dx''' x'''^2 \int_{0}^{2\pi} d\theta''' \int_{0}^{1/2} dx''' x'''^2 \int_{0}^{2\pi} d\theta''' \cos \theta''$$

$$\times \cos \left[\theta''' + (\theta'' - \theta'') \right] (x''^2 + x'''^2 - 2x''' x'''' \cos \theta''')^{5/6} \qquad (59)$$

In order to obtain the final form of Eq. (59), we have replaced $\theta^{\text{\tiny m}}$ by $\theta^{\text{\tiny m}}\!+\!\theta^{\text{\tiny m}}$ (treating $\theta^{\text{\tiny m}}$ as a constant for the $\theta^{\text{\tiny m}}\!-\!integration$), and then adjusted the limits of the new $\theta^{\text{\tiny m}}\!-\!integration$ from $\theta^{\text{\tiny m}}\!\leftrightarrow 2\pi+\theta^{\text{\tiny m}}$ to $0\leftrightarrow 2\pi$. Now we can rewrite $\cos\left[\theta^{\text{\tiny m}}\!+\!(\theta^{\text{\tiny m}}\!-\!\theta')\right]$ as

$$\cos \left[\theta^{m} + (\theta^{m} - \theta')\right] = \cos \theta^{m} \cos (\theta^{m} - \theta') + \sin \theta^{m} \sin (\theta^{m} - \theta')$$
, (60)

and using the same arguments as were used to develop Eq. (57) from Eq. (55), i.e., dropping the $\sin \theta^m$ dependence because it leads to an odd integrand in θ^m , we obtain

$$T_{9} = -3.44 \left(\frac{1}{64} \pi\right)^{-2} (r/D) (r'/D) \int_{0}^{1/2} dx'' x''^{2} \int_{0}^{2\pi} d\theta'' \cos \theta'' \cos (\theta'' - \theta')$$

$$\times \left\{ \int_{0}^{1/2} dx''' x'''^{2} \int_{0}^{2\pi} d\theta''' \cos \theta''' (x''^{2} + x'''^{2} - 2x''' x'''' \cos \theta''')^{5/8} \right\}. (61)$$

First of all, we note the close relationship between the quantity in the curly brackets in Eq. (61) to the integration defining \mathfrak{G}_3 in Eq. (46). Second, we

note that the only θ'' -dependence in Eq. (61) is $\cos \theta'' \cos (\theta'' - \theta')$. Since

$$\int_{0}^{2\pi} d\theta'' \cos \theta'' \cos (\theta'' - \theta') = \pi \cos \theta' , \qquad (62)$$

we see that Eq. (61) can be rewritten as

$$T_9 = -(D/r_0)^{5/3} (r/D)(r'/D) \cos \theta' G_4$$
 (63)

With all of these results in hand, we can now rewrite $\,C_{\!\phi}\,$ as given by Eq. (33) in the form

$$C_{\varphi}(|\vec{r} - \vec{r}'|; D) = C_{\varphi}(r, r', \theta'; D)$$

$$= (D/r_0)^{\rho/3} \{-G_0(|\vec{r}/D) - (\vec{r}'/D)|\} + G_1(r/D)$$

$$+ G_1(r'/D) - G_2 + (r'/D) \cos \theta' G_3(r/D)$$

$$+ (r/D) \cos \theta' G_3(r'/D) - (r/D)(r'/D) \cos \theta' G_4\}. (64)$$

In writing Eq. (64), we have taken the liberty of introducing the notation $C_{\phi}(r, r', \theta'; D)$ to make it explicit that the dependence on $|\vec{r} - \vec{r}'|$ can just as well be considered a dependence on r, r', and θ' . We note that although θ' has been defined in terms of the \vec{r}' , physically, and in our future work, it can be considered as simply being the angle between \vec{r} and \vec{r}' , if we choose.

The first thing we wish to note in considering Eq. (64) is the fact that it is indeed a function of \vec{r}/D and \vec{r}'/D multiplied by $(D/r_0)^{5/3}$, thus justifying the assertion that led us to write down Eq. (21). In fact, comparing Eq.'s (21) and (64), we see that the value of $\mathfrak{C}(|\vec{x}-\vec{x}'|)$ can be written as

$$\mathfrak{C}(\left|\overrightarrow{x}-\overrightarrow{x}'\right|) = \mathfrak{K}(x, x', \theta')$$

$$= -\mathfrak{G}_0\left(\left|\overrightarrow{x}-\overrightarrow{x}'\right|\right) + \mathfrak{G}_1(x) + \mathfrak{G}_1(x') - \mathfrak{G}_2$$

$$+ x' \cos \theta' \mathfrak{G}_3(x) + x \cos \theta' \mathfrak{G}_3(x') - x x' \cos \theta' \mathfrak{G}_4. (65)$$

With this expression in hand for \mathfrak{F} , with the \mathfrak{G} -functions defined by Eq.'s (43) to (47), we can now turn our attention back to the problem of solving the integral equation for the eigenvalues $\mathfrak{R}_{\mathbf{n}}^{2}$ and the eigenfunctions $\mathfrak{F}_{\mathbf{n}}(\vec{\mathbf{x}})$, in accordance with Eq. (28).

Integral Equation Reduction

As indicated at the start of this paper, we shall not attempt here to solve the Karhunen-Loeve homogeneous integral equation, the pertinent form of which is given by Eq. (28). Ultimately we intend to solve this equation using numerical techniques, in particular the Givens-Householder method. At this point, however, we are interested in numerical procedures that will simplify this equation. We note that the basic integral equation is two-dimensional and that as a consequence, the size of the matrix required to obtain any reasonable resolution over the aperture for our eigenfunctions will be unreasonably large. We propose to avoid this problem by introducing a separation of variables in the eigenfunction.

We postulate that the eigenfunction $\mathfrak{F}_n(\vec{x})$ can be separated into a radial dependence and an azimuthal dependence, and further postulate that the azimuthal dependence can be written in the form $\exp(i q \theta)$, where q=0, ± 1 , ± 2 , . . . is a "quantum number" for a set of solutions corresponding to a subset of the n "quantum numbers." The radial dependence would have its own set of "quantum numbers", p=1, 2, 3, . . . , with each combination (p,q) corresponding to an element in the set that was "counted" by n . Thus we would write

$$\mathfrak{J}_{\mathbf{n}}(\vec{\mathbf{x}}) \equiv \mathfrak{R}_{\mathbf{p}}^{\mathbf{q}}(\mathbf{x}) \exp(i \mathbf{q} \theta) ,$$
 (66)

where the superscript q over the radial dependence term, i.e., the function \Re , is used to indicate that for each value of q, we may expect to obtain a different set of radial dependence functions. The eigenvalue would now be written as $\Re^2_{\mathbf{p},\mathbf{q}}$ in place of $\Re^2_{\mathbf{p}}$.

To validate our hypothesis concerning the validity of Eq. (66), we shall substitute Eq. (66) for $\mathfrak{F}_n(\vec{x})$ into the Karhunen-Loeve homogeneous integral equation given by Eq. (28), and show that it leads to self-consistency in the sense that the form of $\mathfrak{F}_n(\vec{x})$ will be found to have the form given by Eq. (66). While this is not a rigorous proof of the validity of the separation of variables, our manipulations will define the key steps required to develop such a rigorous proof. We shall not concern ourselves further with the matter of a rigorous proof.

If we substitute Eq. (66) for $\mathfrak{J}_{\mathbf{n}}(\vec{x})$ into Eq. (28), we obtain

$$\int d\vec{x}' W(\vec{x}', 1) \, g(|\vec{x} - \vec{x}'|) \, g_n(\vec{x}')$$

$$= \int_0^{1/2} dx' \, x' \int d\theta' \, g(x, x', \theta' - \theta) \, g_p^q(x') \, \exp(i \, q \, \theta') \quad , \tag{67}$$

where here we have chosen an arbitrary angular reference point so that we can define angles θ and θ' associated with \vec{x} and \vec{x}' , respectively, rather than merely having θ' defined as the angle between \vec{x} and \vec{x}' . Then in writing G, we took note of the fact that the θ' -dependence indicated in Eq. (65) was in this case a dependence on $\theta' - \theta$. Now if we replace θ' in Eq. (67) with $\theta' + \theta$ and then readjust the limits of the θ' -integration from $\theta \leftrightarrow 2\pi + \theta$ to $0 \leftrightarrow 2\pi$, we obtain

$$\int d\vec{x}' W(\vec{x}', 1) \mathcal{E}(|\vec{x} - \vec{x}'|) \mathcal{B}_n(\vec{x}')$$

$$= \exp(i q \theta) \int_0^{1/2} dx' \mathcal{R}_q(x, x') \mathcal{R}_p^q(x') , \qquad (68)$$

where we have used Ra to mean

$$\Re_{q}(x, x') = x' \int_{0}^{2\pi} d\theta' \, \mathfrak{T}(x, x', \theta') \, \exp\left(i \, q \, \theta'\right) \quad . \tag{69}$$

Now combining Eq. 's (68) and (28), we see that

$$\mathfrak{J}_{n}(\vec{\mathbf{x}}) = \exp\left(i \ \mathbf{q} \ \mathbf{\theta}\right) \left\{ \mathfrak{B}_{n}^{-2} \int_{0}^{1/2} d\mathbf{x}' \ \mathbf{R}_{q}(\mathbf{x}, \mathbf{x}') \, \mathfrak{R}_{p}^{q}(\mathbf{x}') \right\}. \tag{70}$$

Comparison of Eq. (70) with Eq. (66) shows that our assumption of Eq. (66) for $\mathfrak{F}_n(\vec{x}')$ leads to self-consistent results for $\mathfrak{F}_n(\vec{x})$. Combining Eq.'s (66) and (70), we obtain the equation

$$\int_{0}^{1/2} dx' \, \Re_{q}(x, x') \, \Re_{p}^{q}(x') = \Re_{p, q}^{2} \, \Re_{p}^{q}(x) \qquad . \tag{71}$$

This is a Karhunen-Loeve integral equation type definition for the radial function, \Re_p^q , with a kernel, \Re_q , which can be different for each value of the "quantum number," q.

Eq. (71) provides us with a basis for calculation of the complete set of eigenvalues, $\mathfrak{B}_{p,q}^2$ (or \mathfrak{B}_n^2 in our original notation), and together with Eq. (66) provides a definition of our two-dimensional eigenfunction, $\mathfrak{R}_p^q(x)$ exp (i q 0) (or $\mathfrak{F}_n(\vec{x})$ in our original notation). The set of kernels, $\{\mathfrak{R}_q\}$, can be obtained by substituting Eq. (65) into Eq. (69). We get

$$R_{0}(x, x') = -x' \int_{0}^{2\pi} d\theta' G_{0}([x^{2} + x'^{2} - 2x x' \cos \theta']^{1/2} + 2\pi x' [G_{1}(x) + G_{1}(x') - G_{2}] , \qquad (72)$$

$$R_{\pm 1}(x, x') = -x' \int_{0}^{2\pi} d\theta' \, G_{0}([x^{2} + x'^{2} - 2x \, x' \cos \theta']^{1/2}) \cos (\theta')$$

+
$$\pi \times' [\times' \otimes_3 (x) + \times \otimes_3 (x') - \times \times' \otimes_4]$$
, (73)

and for the magnitude of q greater than one

$$R_{q}(x, x') = -x' \int_{0}^{2\pi} d\theta' \, \Theta_{0} \, ([x^{2} + x'^{2} - 2x \, x' \cos \theta']^{1/2}) \cos (q + t')$$

$$for \, q = \pm 2, \, \pm 3, \, \pm 4, \, \dots$$
(74)

In obtaining these results, we have made use of the fact that exp (i q θ') can be written as $\cos \theta' + i \sin \theta'$, and noted that the $\sin \theta'$ leads to integrands odd in θ' so that their value after integration of $0 \leftrightarrow 2\pi$ vanishes.

The combination of Eq.'s (26), (27), (43), (44), (45), (46), (47), (66), (71), (72), (73), and (74) provides the basis for calculating the eigenvalues and eigenfunctions we shall need to determine the probability of an accidental occurrence of a low energy wavefront distortion condition so that the CENSORING system will be able to produce a near diffraction-limited image. In the next section, we take up the problem of calculating this probability, given the set of eigenvalues. As noted before, we leave the problem of numerically evaluating the eigenvalues and eigenfunctions for treatment in a subsequent paper.

Probability Formulation

The key to the evaluation of the probabilities associated with a CENSORING system's performance is to recognize that this is essentially equivalent to a study of the probability distribution of the effective meansquare wavefront distortion over the aperture. The term "effective" as used here refers to the fact that we are only interested in wavefront variations excluding tilt and average phase variations -- i.e., the effective wavefront distortion is to be calculated from $\phi(\vec{r};D)$ and not from $\phi(\vec{r})$. We write the mean-square wavefront distortion over the aperture as

$$\Delta^{2} = \left(\frac{1}{4} \pi D^{2}\right)^{-1} \int d\vec{r} W(\vec{r}; D) \left| \varphi(\vec{r}; D) \right|^{2} \qquad (75)$$

It should be recognized that the term "mean" in reference to Δ^2 refers to an average over the aperture and not to an ensemble average. Thus just as the effective wavefront distortion, ϕ , is a random function, we see from Eq. (75) that Δ^2 is a random variable. If, at some instant, Δ^2 is small enough, then we may expect nearly diffraction-limited quality for an image formed at that instant. The problem we face in calculating CENSORING system performance is one of calculating the probability that Δ^2 will be small enough. We set as a nominal threshold the requirement that $\Delta^2 < \Delta_1^2$ radian-square as the dividing line between good and poor images. Our problem is to calculate the probability of Δ^2 being less than Δ_1^2 , this being the probability that the CENSORING system will see good enough conditions to allow an image to be formed.

If we substitute Eq. (6) into Eq. (75), we get

$$\Delta^{2} = (\frac{1}{4} \pi D^{2})^{-1} \sum_{n, n'} \beta_{n} * \beta_{n'} \int d\vec{r} W(\vec{r}, D) f_{n} * (\vec{r}) f_{n'}(\vec{r}) . \qquad (76)$$

Now making use of the orthonormality of f_n , as defined in Eq. (5), we can reduce Eq. (76) to the form

$$\Delta^{2} = (\frac{1}{4} \pi D^{2})^{-1} \sum_{n} \beta_{n}^{*} \beta_{n}$$

$$= \sum_{n} [\beta_{n}/(\frac{1}{4} \pi D^{2})^{1/2}]^{*} [\beta_{n}/(\frac{1}{4} \pi D^{2})] \qquad (77)$$

Thus the mean-square wavefront distortion is seen to be the sum of the square of a set of gaussian random variables, $\beta_n/(\frac{1}{4} \pi D^2)^{1/2}$. We recall that according to Eq. (8), the random variable β_n has a variance given by the eigenvalue $B_n^2(D)$, so that the variance of $\beta_n/(\frac{1}{4} \pi D^2)^{1/2}$ can be written, using Eq. (27), in the form

$$Var \left[\beta_{n} / \left(\frac{1}{4} \pi D^{2} \right)^{1/2} \right] = \left(\frac{1}{4} \pi D^{2} \right)^{-1} B_{n}^{2} (D)$$

$$= \left(D / r_{0} \right)^{5/3} \left(4 / \pi \right) \mathfrak{B}_{n}^{2} . \tag{78}$$

For convenience, we shall denote this variance by $\sigma_{\mathbf{p},\mathbf{q}}^{2}$ (or $\sigma_{\mathbf{p},\mathbf{q}}^{2}$) as appropriate.

$$\sigma_n^2 = \text{Var} \left[\beta_n / (\frac{1}{4} \pi D^2)^{1/2}\right]$$
 (79)

If we can compute the eigenvalues, \mathfrak{P}_n^2 (or $\mathfrak{P}_{p,q}^2$) for the dedimensionalized Karhunen-Loeve Homogeneous Integral equation of Eq. (28) [or of Eq. (71)], then we can immediately write

$$\sigma_n^2 = (D/r_0)^{5/3} (4/\pi) \vartheta_n^2 ,$$
 (80)

or

$$\sigma_{p,q}^2 = (D/r_0)^{5/3} (4/\pi) \Re_{p,q}^2$$
 (80')

It now follows that the probability of CENSORING system at any instant seeing low enough distortion to allow an image to be formed can be written as

$$P_{Cengor}$$
 = Prob (CENSORING System Forming an Image)
= Prob ($\Delta^2 \le \Delta_T^2$) . (81)

Since the random variables $\,\beta_{n}\,\,$ are independent and gaussian distributed with variance $\,\sigma_{n}^{\,\,2}\,\,$, it follows that

$$P_{consor} = \prod_{n=1}^{\infty} (2\pi \sigma_n^2)^{-1/2} \int dx_n \exp(-\frac{1}{2} x_n^2/\sigma_n^2) , \qquad (82)$$

where the limits on the integration are to be understood as a composite limit on the product, or rather on the n-tuple multiple integral. The limit

corresponds to

Limit
$$\equiv \left(\sum_{n=1}^{\infty} x_n^2 \le \Delta_{\tau}^2\right)$$
 (83)

Eq.'s (82) and (83), in concept at least, provide a basis for the calculation of P_{Conpor} , i.e., the probability of a CENSORING system image being formed. However, because of the infinite limits on n in these two equations, no practical calculations can be performed.

To provide a practical basis for carrying out the calculation of P_{consor} , we need to truncate the series. To do this, we first note that in accordance with results which we have previously obtained elsewhere, we know that the ensemble average value of Δ^2 can be written as

$$\langle \Delta^2 \rangle = 0.1345 \, (D/r_0)^{6/3} \quad . \tag{84}$$

Now if we assume that the eigenfunctions are arranged in such an order that σ_n^2 is monotonically decreasing with n, then we know that if N is large enough so that

$$\sum_{n=1}^{N} \sigma_n^2 = \langle \Delta^2 \rangle - \varepsilon \Delta_{\tau_n} \qquad (85)$$

then if ε has been chosen to be a small enough quantity, we may consider N, rather than ∞ , to be the practical upper limit on the n dependencies in Eq. 's (82) and (83). As a practical matter, we would replace Δ_{τ}^2 in Eq. (83) with $\Delta_{\tau}^2(1-\varepsilon)$.

This has the effect of saying that above some value of n (namely n=N), we are not particularly concerned with the exact amount of wavefront distortion introduced by each degree of freedom. The exact value

of those β_n 's does not concern us since we know that the corresponding variances, σ_n^2 , are so small that the values of the β_n 's will be tolerably small. We expect the contribution of all those higher order terms, which we are suppressing, to the mean-square wavefront distortion to only be of the order of $\epsilon\Delta_{\tau}^2$, which, by our choice of ϵ , we have made sure is tolerably small.

In an actual calculation of P_{consor} , we would first compute an ordered series of eigenvalues \mathfrak{B}_n^2 . Then using Eq. (80), we would compute the variances, σ_n^2 . By applying Eq. (85), we could then determine the truncation level, N. At that point, our calculation would reduce to the evaluation of the integral

$$P_{consor} = \iint ... \int dx_1 dx_2... dx_N \prod_{n=1}^{N} (2^{\pi} \sigma_n^2)^{-1/2} \exp(-\frac{1}{2} x_n^2/\sigma_n^2), \quad (86)$$

where

Limit
$$\equiv \left(\sum_{n=1}^{N} x_n^2 \leq \Delta_{\tau}^2 (1-\epsilon)\right)$$
 (87)

Our problem is thus reduced to the numerical evaluation of the integral in Eq. (86), having first solved the Karhunen-Loeve homogeneous integral equation for the eigenvalues. Neither is a trivial numerical task, but in practice can be expected to be rather straightforward, if somewhat ponderous in terms of required computer effort. These tasks are taken up in Part II of this report.

PART II

Numerical Evaluation of Probabilities

Governing the Performance of a

CENSORING System

Introduction

In Part I of this report, a formal basis was developed for the analysis of the expected performance of a CENSORING system. The basic quantity of interest was identified as the probability that at any instant of time, the wavefront distortion over the entrance aperture would have an rms deviation from a plane of less than one radian. (The term "rms" is used here in the sense of an average over the aperture.) Because the CENSORING system is designed to form a short exposure image, the wavefront deviation is to be measured relative to the optimally chosen tilted plane, i.e., that plane whose tilt is such as to minimize the rms deviation.

It was shown that the probability of interest could be calculated in terms of a multi-dimensional gaussian distribution in a hyper-space, where the hyper-space was defined in terms of a set of functions which could be used to decompose a sample of the randomly distorted wavefront taken over the aperture into a set of statistically independent components. Because the wavefront distortion is a gaussian random process, 1 and because the magnitude of each of these independent random components is obtained by a linear process from the random wavefront distortion, it follows that each component is a gaussian random variable. The random amplitude of each component represents one of the dimensions in the hyperspace, and the probability of interest is the probability that all of the random variables will take on small enough values at some instant of time.

Because the functions used for the decomposition of the wavefront are a set of functions that are orthonormal over the space defined by the CENSORING system's aperture, it follows that the mean square deviation of the wavefront at any instant (with the mean taken as an average over the aperture) is just equal to the sum of the square of the component amplitudes

in the wavefront decomposition. This means that the probability of the rms wavefront distortion being less than or equal to one-radian is just the probability that the random components will define a point in the hyperspace that lies within a hyper-sphere of one-radian radius and centered at the origin. Since the components each obey an independent gaussian distribution, the problem of evaluating the probability of interest can be seen to reduce to a multi-dimensional integral within a unit hyper-sphere of a set of gaussian distributions. All we need in order to be able to carry out this evaluation is information on the variances to be associated with each element of the set of gaussian distribution. In the previous work, it was shown that these variances could be obtained by solving the Karhunen-Loève integral equation associated with the wavefront distortion statistics, and that they were proportional to the eigenvalues of that equation.

The single wavefront distortion statistic required for this problem is the wave structure function, $\mathcal{F}(r)$, where

$$\mathcal{L}(r) = 6.88 (r/r_0)^{5/3} (1)$$

The orthonormal function set $\{f_n(\vec{r})\}$ and the associated set of variances, $\{\sigma_n^2\}$, which appropriately decompose the distorted wavefront and which define the gaussian probability distributions of interest in our hyper-space integration have been shown to correspond to the eigenfunctions and to be proportional to the eigenvalues of a Karhunen-Loève integral equation utilizing a function of $\mathcal{P}(r)$ as the kernel of the integral. The relationship between the set of variances of interest and the set of eigenvalues, $\{B_n^2\}$, is given by the equation

$$\sigma_n^2 = B_n^2 / (\frac{1}{4} \pi D^2)$$
 (2)

In seeking a solution of this integral equation, it has been shown that the set $\{f_n(\vec{r})\}$ can be decomposed in subsets by separation of variables

into polar coordinates, i.e.,

$$\vec{r} \equiv (r, \theta) \qquad . \tag{3}$$

It has been shown that we can write

$$f_n(\vec{r}) \equiv R_p^q(r) \exp(i q \theta)$$
, $q = -\infty, ... -2, -1, 0, 1, 2, ... +\infty$
 $p = 1, 2, 3, ... +\infty$, (4)

where the function $R_p^q(r)$ represents a set, on p, of functions that satisfy a homogeneous integral equation with a kernel that is different for each value of q. (Actually the kernel for q and -q are identical.) The eigenvalues of this integral equation, $B_{p,q}^{2}$ can be equated with the eigenvalues of the original integral equation, i.e.,

$$B_{p,q}^2 = B_n^2 , \qquad (5)$$

and equivalently, the variance of interest can be written in the form $\sigma_{p,q}^{-2}$, where

$$\sigma_{p,q}^2 = \sigma_n^2 \qquad . \tag{6}$$

The integral equation defining $R_{p,q}(r)$ and $B_{p,q}^{-2}$ involves the aperture diameter of the CENSORING system, D, and the basic wavefront distortion turbulence parameter, r_0 . It is convenient to cast the integral equation in a form which is independent of these two parameters so that a single set of numerical solutions to the integral equation can be applied for all possible values of D and r_0 . It has been shown that if we define the eigenfunctions, $\Re_p^q(x)$ and $\Re_{p,q}^{-2}$ by the integral equation

$$\int_{0}^{1/2} dx' \Re_{q}(x, x') = \Re_{p, q} \Re_{p}^{q}(x) , \qquad (7)$$

then

$$B_n^2 = B_{p,q}^2 = D^{11/3} r_0^{-6/3} B_{p,q}^2$$
, (8)

and

$$f_{\mathbf{n}}(\mathbf{r}) = R_{\mathbf{p}}^{\mathbf{q}}(\mathbf{r}) \exp (i \mathbf{q} \theta) = \Re_{\mathbf{p}}^{\mathbf{q}}(\mathbf{r}/\mathbf{D}) \exp (i \mathbf{q} \theta)$$
 (9)

If we define $\tilde{R}_a(x, x')$ as

$$\tilde{R}_{q}(x, x') = -x' \int_{0}^{2\pi} d\theta' \mathcal{G}_{0}([x^{2} + x'^{2} - 2xx' \cos \theta']^{1/2}) \cos (q\theta'), (10)$$

then the kernel for the integral equation of Eq. (7) can be written as

$$\Re_{q}(x, x') = \widetilde{\Re}_{q}(x, x')$$
 if $q = \pm 2, \pm 3, \pm 4, \ldots$, (11)

$$\Re_{1}(\mathbf{x}, \mathbf{x}') = \widetilde{\Re}_{1}(\mathbf{x}, \mathbf{x}') + \pi \mathbf{x}' [\mathbf{x}' \mathbf{G}_{3}(\mathbf{x}) + \mathbf{x} \mathbf{G}_{3}(\mathbf{x}') - \mathbf{x} \mathbf{x}' \mathbf{G}_{4}] , \qquad (12)$$

$$\Re_{0}(\mathbf{x}, \mathbf{x}') = \widetilde{\Re}_{0}(\mathbf{x}, \mathbf{x}') + 2\pi \mathbf{x}' [\mathfrak{G}_{1}(\mathbf{x}) + \mathfrak{G}_{1}(\mathbf{x}') - \mathfrak{G}_{2}] \qquad (13)$$

The G-functions are defined as

$$\mathfrak{G}_{b}(x) = 3.44 x^{5/3}$$
 , (14)

$$\mathfrak{G}_{1}(x) = 3.44 \left(\frac{1}{2}\pi\right)^{-1} \int_{0}^{1/2} dx'' x'' \int_{0}^{2\pi} d\theta'' (x^{2} + x''^{2} - 2xx'' \cos \theta'')^{5/6}, (15)$$

$$\mathfrak{G}_{2} = 8 \int_{0}^{1/2} dx'' x'' \, \mathfrak{C}_{1}(x'') , \qquad (16)$$

$$G_3(x) = 3.44 \left(\frac{1}{2}\pi\right)^{-1} \int_{0}^{1/2} dx \, x \, x^{2} \int_{0}^{2\pi} d\theta \, \cos \theta \, (x^2 + x^{2}) d\theta \, \cos \theta \, (x^2 + x^{2}) d\theta \, \cos \theta \, (x^2 + x^{2}) d\theta \, \cos \theta \, \cos \theta$$

$$\mathfrak{G}_{4} = 64 \int_{0}^{1/2} dx \, \, x \, \, x \, \, \mathfrak{G}_{3}(x'') \qquad . \tag{18}$$

We note in passing that if we had been interested in a system in which, even during a short exposure, wavefront tilt had to be considered a portion of the wavefront distortion (which it does not have to be for normal short exposure purposes), then the only change in the above results would have been to modify Eq. (12) to the form

$$R_1(\mathbf{x}, \mathbf{x}') = \widetilde{R}_1(\mathbf{x}, \mathbf{x}') \qquad . \tag{12'}$$

Our basic problem is to solve the integral Eq. (7) for all of the eigenvalues, $\{\mathfrak{P}_{\mathfrak{p},\mathfrak{q}}^{2}\}$ and then for a particular set of values of D and \mathfrak{r}_{0} obtain the corresponding set of variances $\{\mathfrak{C}_{\mathfrak{p},\mathfrak{q}}^{2}\}$ in accordance with Eq.'s (2) and (8), from which we get

$$\sigma_{p, q}^{2} = \frac{4}{\pi} \left(\frac{D}{r_{0}} \right)^{5/3} \Re_{p, q}^{2} .$$
 (19)

The second half of the problem is the evaluation of the integral in hyperspace which defines the probability that at any instant the random wavefront distortion relative to the optimally chosen tilted plane will have a mean square value, averaged over the aperture, of one radian squared or less. This probability can be written as

$$P_{CENSOR} = \iint_{p,q} (2\pi \, \sigma_{p,q}^2)^{-1/2} \int_{sph} dx_{p,q} \, \exp\left(-\frac{1}{2} \, x_{p,q}^2/\sigma_{p,q}^2\right) , (20)$$

where the limits on the integral correspond to a hyper-sphere for which

$$\sum_{p,q} x_{p,q}^2 \leq 1 \qquad . \tag{21}$$

In Eq. 's (20) and (21), the product and the summation run over all possible combinations of values of p and q.

In the next section, we take up the problem of casting the integral equation of Eq. (7) in a form suitable for numerical evaluation. In the sections after that, we shall first present the numerical solution technique and results, and then go into the problem of formulating an explicit form for the hyper-space integral of Eq. (20) for various values of D/r_0 . Then we shall move on to consider the evaluation of that hyper-space probability integral. Based on the results of this evaluation, we shall present a general discussion of the expected performance of a CENSORING system.

Integral Equation Numerical Formulation

The numerical solution of the Karhunen-Loève integral equation presented in Eq. (7) can be developed using standard numerical techniques. However, it will greatly simplify our numerical treatment if we first recast that equation into an equivalent form in which the kernel, i.e., $\Re_q(\mathbf{x},\mathbf{x}')$ is replaced by a kernel that manifests symmetry between \mathbf{x} and \mathbf{x}' . [We note that the leading factor of \mathbf{x}' in the right hand side of Eq.'s (10), (12), and (13) destroys the symmetry of $\Re_q(\mathbf{x},\mathbf{x}')$.]

In order to obtain the desired symmetry, we introduce the following symmetrizing functions:

$$\tilde{R}_{q}^{s}(x, x') = -(x x')^{1/2} \int_{0}^{2\pi} d\theta' \mathcal{G}_{0}([x^{2} + x'^{2} - 2x x' \cos \theta']^{1/2}) \cos (q \theta'), \quad (22)$$

$$R_0^s(x, x') = \tilde{R}_0^s(x, x') + 2\pi (x x')^{1/2} [G_1(x) + G_1(x') - G_2]$$
, (23)

$$R_1^{s}(x, x') = \widetilde{R}_1^{s}(x, x') + \pi (x x')^{1/2} [G_3(x) + G_3(x') - G_4]$$
, (24)

$$R_{-1}^{s}(x, x') = R_{1}^{s}(x, x')$$
 (25)

$${}^{5}\mathfrak{R}_{p}{}^{q}(x) = x^{1/2}\mathfrak{R}_{p}{}^{q}(x)$$
 (26)

Now it follows from direct substitution that Eq. (7) can be rewritten as

$$\int_{0}^{1/2} dx' \, \Re_{q}^{s}(x, x') \, {}^{s}\Re_{p}^{q}(x') = \Re_{p, q}^{2} \, {}^{s}\Re_{p}^{q}(x) , \qquad (27)$$

which integral equation has a symmetric kernel and can be solved using straightforward numerical techniques. The eigenvalues of Eq. (27) are identical to those of Eq. (7), and the eigenfunctions of Eq. (7), i.e., $\Re_p^q(x)$ can be obtained from the eigenfunctions of Eq. (27), i.e., $\Im_p^q(x)$ by use of Eq. (26).

To obtain the eigenvalues and eigenfunctions of Eq. (27), it is necessary to transform the integration into a summation, thereby obtaining a homogeneous set of simultaneous equations with a determinant that determines the eigenvalues and eigenfunctions. In order to replace the integration by a summation, we subdivide the range x=0 to 0.5 into 20 sections and consider the values of x at the midpoint of each section, which we denote by x_t for t=1 to 20. We can then make the replacement

$$\int_{0}^{1/2} d\mathbf{x}' \, \mathbf{R}_{q}^{s} (\mathbf{x}, \mathbf{x}')^{s} \, \mathfrak{R}_{p}^{q} (\mathbf{x}') \Rightarrow \sum_{i=1}^{20} K_{q}^{s} (i, i')^{s} \, \mathbf{R}_{p}^{q} (i') \qquad , \tag{28}$$

which allows us to rewrite Eq. (27) as

$$\sum_{i=1}^{20} K_{q}^{s}(i,i')^{s} R_{p}^{q}(i') = \Re_{p,q}^{2s} R_{p}^{q}(i) , \qquad (29)$$

where

$$R_{n}^{q}(i) \equiv {}^{5}\mathfrak{R}_{n}^{q}(\mathbf{x}_{i}) , \qquad (30)$$

and

$$K_a^{s}(i,i') = (\frac{1}{2}/20) R_a^{s}(x_t,x_t)$$
 (31)

Eq. (29) presents us with the straightforward problem of obtaining the eigenvalues and eigenfunctions of the matrix K_q^s (i.i'). This is a

standard problem in numerical analysis. The 20 x 20 size of the matrix means that we are dealing with a rather small problem as such computations go on a large high-speed digital computer. The mathematical procedure that we shall use is based on the well-known Givens-Householder algorithm² and associated procedures. (The details of the actual computation will not be discussed, as the computer program utilized a proprietary CDC subroutine for this task.)

Computationally, the evaluation of K_q^s (i, i') matrix elements represented almost as large a task as the determination of the eigenvalues and eigenfunctions of this matrix. The pertinent expressions can be written as

$$K_0^{5}(i, i') = 6.88 (.025)(x_i x_i)^{1/2} \left\{ -\int_{0}^{\pi} d\theta (x_i^2 + x_i)^2 - 2x_i x_i, \cos \theta)^{5/6} + 8 \int_{0}^{1/2} du u \int_{0}^{\pi} d\theta [(x_i^2 + u^2 - 2x_i u \cos \theta)^{5/6} + (x_i^2 + u^2 - 2x_i u \cos \theta)^{5/6} + (x_i^2 + u^2 - 2x_i u \cos \theta)^{5/6} + (x_i^2 + u^2 - 2x_i u \cos \theta)^{5/6} \right\} - 64 \int_{0}^{1/2} du u \int_{0}^{1/2} dv v \int_{0}^{\pi} d\theta (u^2 + v^2 - 2u v \cos \theta)^{5/6} \right\} , \quad (32)$$

$$K_{\pm i}^{5}(i, i') = 6.88 (.025)(x_{i} x_{i})^{1/2} \left\{ -\int_{0}^{\pi} d\theta (x_{i}^{2} + x_{i})^{2} - 2x_{i} x_{i}, \cos \theta \right\}^{6/6} \cos \theta$$

$$+ 64 \int_{0}^{1/2} du u^{2} \int_{0}^{\pi} d\theta [x_{i}(x_{i})^{2} + u^{2} - 2x_{i}, u \cos \theta]^{6/6} + x_{i}, (x_{i}^{2} + u^{2} - 2x_{i})^{2} u \cos \theta + u^{2} - 2x_{i} u \cos \theta]^{6/6} \right\} \cos \theta$$

$$- 4096 x_{i} x_{i} \int_{0}^{1/2} du u^{2} \int_{0}^{1/2} dv v^{2} \int_{0}^{\pi} d\theta (u^{2} + v^{2} - 2uv \cos \theta)^{6/6} \cos \theta \right\}, (33)$$

and

$$K_q^{s}(i, i') = -6.88 (.025)(x_1 x_1)^{1/2} \int_0^{\pi} d\theta (x_1^2 + x_1^2 - 2x_1 x_1, \cos \theta)^{5/6} \cos (q\theta),$$
for $|q| > 1$. (34)

We have utilized the trapezoidal rule to carry out the θ -integrations in Eq.'s (32), (33), and (34) and Romberg interpolation³, and have used a 10-point Gaussian quadrature to evaluate the u- and v-integrations.

For a value of q much greater than two (2), there is a potential accuracy problem in the evaluation of the integral in Eq. (34). This is most easily studied if we split the θ -integration in Eq. (34) in such a way that each region of integration goes over only one-half cycle of the oscillating function $\cos (q\theta)$. This gives us

$$\int_{0}^{\pi} d\theta (x_{1}^{2} + x_{1}^{2} - 2x_{1}^{2} x_{1}, \cos \theta)^{6/6} \cos (q\theta)$$

$$= \sum_{k=1}^{q} \int_{0}^{k \pi/q} d\theta (x_{1}^{2} + x_{1}^{2} - 2x_{1}^{2} x_{1}, \cos \theta)^{6/6} \cos (q\theta). (35)$$

The factor $(x_1^2 + x_1^2 - 2x_1 x_1, \cos \theta)^{5/8}$ is a positive monotonically-increasing function of θ in Eq. (35), from which we see that the right-hand-side of Eq. (35) consists of the sum of a set of alternating sign terms. When there is a large difference between x_1 and x_1 , we note that $(x_1^2 + x_1^2 - 2x_1 x_1^2, \cos \theta)^{5/6}$ is a very weak function of θ , so that the terms being summed are nearly equal in magnitude, but of alternating sign -- a situation that can seriously stress the accuracy of the computed results.

To avoid this accuracy problem, we can expand the integrand in Eq. (34) in the form

$$\int_{0}^{\pi} d\theta \left(x_{1}^{2} + x_{1}^{2} + x$$

where

$$S = x_1^2 + x_1^2, (37)$$

$$\epsilon = 2 x_1 x_1 / S \tag{38}$$

$$P_{n} = \frac{5}{6} \left(\frac{5}{6} - 1 \right) \left(\frac{5}{6} - 2 \right) \cdot \cdot \cdot \left(\frac{5}{6} - n + 1 \right) \qquad (39)$$

The integral in the final form of the right-hand-side of Eq. (36) can be shown to have the value4

$$\int_{0}^{\pi} d\theta \cos^{n}\theta \cos (q\theta) = \begin{cases} 0 & \text{if } n < q \\ 0 & \text{if } n+q = \text{odd} \\ \frac{\pi n!}{2^{2} \left(\frac{n+q}{2}\right)! \left(\frac{n-q}{2}\right)!} & \text{if } n+q = \text{even} \end{cases}$$

so that Eq. (36) may be rewritten as

$$= \pi S^{5/6} \sum_{n=0}^{\infty} \frac{P_{q+2n} (\epsilon/2)^{q+2n}}{(q+m)! m!}.$$
 (41)

For $\varepsilon < .5$, which corresponds to x_1 and x_2 , significantly different in size. Eq. (41) is fairly rapidly convergent and a 40-term summation yields sufficient accuracy. For $\varepsilon > .5$, x_1 and x_2 , are sufficiently close in value that $(x_1^2 + x_1^2 - 2x_1x_1, \cos\theta)^{5/8}$ is a significant function of θ , and the evaluation of the integral in Eq. (34) can proceed by straightforward numerical quadrature without excessive loss of accuracy. For q greater than two (2), we have evaluated the θ -integration in Eq. (34) using either ordinary numerical integration techniques if $2x_1x_1/(x_1^2 + x_1^2)$ is greater than one-half, or used Eq. (41) for values less than one-half.

With these numerical techniques, the K_q^3 (i, i') matrices were evaluated for q=0 to 41. These results are listed in Table I, giving he matrix in upper right triangular form. With these matrices, we were then able to carry out a determination of the eigenfunctions and eigenvalues. These are listed in Table II. The eigenfunctions as listed here have been restored to their unsymmetrized form, i.e., $\Re_p^q(x)$ instead of $\Re_p^q(x)$ by making use of Eq. (26).

The set of all eigenvalues were rank-ordered without regard to p- and q-values. In this procedure, we counted each eigenvalue twice if its q-value was not zero, since it then applied to both q and -q. This set of eigenvalues is listed in Table III. We note that the leading eigenvalue for each value of q for $q \ge 4$ appear in order according to the value of q. Since we only worked with $q \le 41$, and since the leading eigenvalue for q = 41 is the 569, we can probably consider the list of eigenvalues complete up to the 569, or thereabout. The sum of all the eigenvalues listed, of which there are 1660, is 0.105127. This is in good agreement with the value of 0.1056 expected for the total of all eigenvalues, as derived from an earlier work which considered the expected mean square wavefront distortion. We note that the cumulative sum at the 569^{th} eigenvalue is 0.104708/0.10527 = 99.60% of the total of the eigenvalues listed, and 0.104708/0.1056 = 99.15% of the total of all the eigenvalues.

With this list of eigenvalues, it is possible to proceed immediately to the imaging probability evaluation aspect of the problem. This we take up in the next section. Before turning to that, however, we first note that because of the ease with which we could adapt our mathematics to the case in which tilt is considered to be a significant wavefront distortion, we have carried out such calculations. As noted previously, this involves nothing more complex than using Eq. (34) in place of Eq. (33) for q = 1. With

this replacement, we obtain the symmetrized kernel shown in Table Ia, and the eigenvalues and eigenfunctions (with symmetrization removed) shown in Table IIa. It is particularly interesting to note that the first eigenvalue in this case is about $3\frac{1}{2}$ times larger than the sum of all the eigenvalues when tilt effects are suppressed as not contributing to effective wavefront distortion, and that the first eigenfunction appears to be very nearly a simple tilt. We also note that the second and subsequent eigenvalues with tilt distortion allowed (Table IIa) are very nearly equivalent to the first and subsequent eigenvalues when tilt distortion is not allowed (Table II, q=1).

Probability Integrals

Having the list of eigenvalues given in Table III reliable out to the 569th eigenvalue, and thus reliably containing more than 99% of the sum of all eigenvalues, we are now in a position to start the evaluation of the probability integral governing the performance of a CENSORING system. We recall that this integral is given by Eq. (20). Rewriting this to work with the n-notation (overall rank order) of Table III, rather than with the p, q-notation, we write

$$P_{\text{CENSOR}} = \iint_{n} (2\pi \sigma_{n}^{2})^{-1/2} \int_{\text{Sph}} dx_{n} \exp(-\frac{1}{2}x_{n}^{2}/\sigma_{n}^{2}) , \qquad (42)$$

where the "Sph" limit on the n-dimensional integration corresponds to the constraint

$$\sum_{n} x_n^2 \le 1 \quad . \tag{43}$$

In accordance with Eq. (19), with \mathfrak{B}_n denoting the n^{th} eigenvalue in Table III, we have

$$\sigma_n^2 = \frac{4}{\pi} (D/r_0)^{5/3} \mathfrak{R}_n^2 , \qquad (44)$$

for an aperture of diameter $\,D\,$ with wavefront distortion characterized by the length $\,r_{\!0}\,$.

We plan to evaluate the n-dimensional integral in Eq. (42) by Monte Carlo techniques. As an immediate reduction in the magnitude of the problem, we recognize that σ_n^2 decreases rather rapidly with increasing n, and that beyond some cut-off value of n, the expected value of the sum of the square of the random variables is very tightly constrained to the sum of the σ_n^2 , with very little variability. The real variability in the overall sum of the squares comes from the much fewer σ_n^2 is large and decreases rapidly with increasing n.

To establish the cut-off value of n , namely, N_e , we arbitrarily allow the random variables $x_{\rm s}$ beyond the cut-off to have an expected value of 0.1 for the sum of their values squared, i.e.,

$$\langle \int_{N_c}^{\infty} x_n^2 \rangle = \sum_{N_c}^{\infty} \sigma_n^2 = 0.1 \qquad . \tag{45}$$

Since we know that

$$\sum_{1}^{\infty} \sigma_{n}^{2} = 0.1345 \, (D/r_{0})^{5/3} \qquad , \tag{46}$$

then we can obtain N_c from the running cumulative value of the eigenvalues in Table III. We write in accordance with Eq. (44)

$$\frac{4}{\pi} \left(D/r_0 \right)^{5/3} \quad \sum_{n=0}^{N_0} \ \mathfrak{B}_n^2 = 0.1345 \left(D/r_0 \right)^{5/3} - 0.1 \quad , \tag{47}$$

from which it follows that

$$\frac{N_c}{\sum_{i}^{N_c}} \mathfrak{B}_{n}^{2} = 0.1056 - 0.07854/(D/r_o)^{5/3} \qquad . \tag{48}$$

We see that to avoid pushing the cut-off, N_e , beyond the reliable limit of our eigenvalue table, i.e., beyond about 569, the largest value of D/r_0 we can use is 14.60 (which we take to be 15).

For each value of D/r_0 , we determine N_e in accordance with Eq. (48) and Table II, and then proceed with the evaluation of the truncated probability integral of Eq.'s (42) and (43), which we now rewrite as

$$P_{\text{SENSOR}} = \iint_{1}^{N_{p}} (2\pi \sigma_{n}^{2})^{-1/2} \int dx_{n} \exp(-\frac{1}{2}x_{n}^{2}/\sigma_{n}^{2}) , \qquad (49)$$

$$x_{n}^{2} \leq 0.9 . (50)$$

In the evaluation of the integral in Eq. (49) by Monte Carlo methods, there are a number of approaches to the random sampling that can be used. First, and most obvious, we consider selecting points uniformly distributed in the hypersphere of Eq. (50), and evaluate the integrand of each point. Unfortunately, the integrand will be very small for most of the points selected, since many values of σ_n , for large n, will be much less than unity. This will give very poor sampling efficiency and an urmanageably large number of samples will be needed to yield even modest accuracy.

A second approach to the random sampling is to select the random points in accordance with the gaussian probability distributions for each dimension inherent in the integrand in Eq. (49). Then the integral would be evaluated by counting a one for each such randomly selected hyper-space

point that satisfied Eq. (50), and zero for each point that did not, and taking the average of these counts of one and zero. Unfortunately, this method also suffers from the problem that an unmanageably large number of sample points are needed to obtain acceptable accuracy in the integral evaluation. In this case, the problem is associated with the smaller values of n, especially for larger values of D/r_0 . The values of σ_n are so large that the gaussian distribution of x_n causes most points selected to lie outside the hypersphere defined by Eq. (50).

To get around the problems of both the first and the second methods of sampling, we used an arbitrarily chosen sampling distribution $\mathfrak{P}_n\left(x_n\right)$ for each of the N_c dimensions of the integral. To compensate for this method of choosing the samples, it is merely necessary to introduce a factor of $\left\lceil\frac{n}{l}\right\rceil\mathfrak{P}_n\left(x_n\right)\right\rceil^{-1}$ into the integrand. We choose $\mathfrak{P}_n\left(x_n\right)$ to match the gaussian distribution with variance σ_n^2 for the larger values of n in the integration, for which σ_n^2 is small and there is, therefore, no tendency to pick values of x_n that are incompatible with Eq. (50). For the smaller values of n, with larger corresponding values of σ_n^2 , we chose a gaussian distribution with a variance σ_0^2 . Here σ_0^2 is the same for all of the dimensions, and significantly less than the corresponding σ_n^2 values. To establish the transition between what we have called the large values of n and the small values of n, we determined a transition value, N_r , which would satisfy the requirement that

$$N_r \sigma_0^2 + \sum_{N_r+1}^{N_s} \sigma_n^2 = 0.9$$
 , (51)

where

$$\sigma_0^2 = \sigma_{N_T}^2 \qquad . \tag{52}$$

The value of N_f is directly obtainable from the eigenvalues of Table III, using Eq. (44).

Using a gaussian sampling distribution with variance $\sigma_0^2 = \sigma_{N_f}^2$ for variables x_1 , x_2 , x_3 , ... x_{N_f} , and variance σ_n^2 for the variables x_{N_f+1} , x_{N_f+2} , ..., x_{N_e} , we have found that Eq. (50) is satisfied between one-third and two-thirds of the time by the hyper-space random vector so chosen. Using this sampling procedure, and noting that now the probability integral of Eq. (49) has the form

$$P_{\text{CENSOR}} = \int_{\substack{\text{random} \\ \text{samples} \\ \{x_n\}_{N_n}}} \frac{(2\pi \ \sigma_n^{\,2})^{-1} \ \exp\left(-\frac{1}{2} \ x_n^{\,2}/\sigma_n^{\,2}\right)}{(2\pi \ \sigma_0^{\,2})^{-1} \ \exp\left(-\frac{1}{2} \ x_n^{\,2}/\sigma_0^{\,2}\right)} \frac{Q\left(\{x_n\}\right)}{\left(\text{number of} \atop \text{samples}\right)}, (53)$$

where

$$Q(\{x_n\}) = \begin{cases} 0 & , & i' \sum_{1}^{N_0} x_n^2 > 0.9 \\ & . & . \end{cases}$$

$$1 & . & . & .$$

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The evaluation of P_{CENSOR} was carried out in accordance with Eq. (53) using samples of 100 points for various values of D/r_0 . By repeating the evaluation a number of times, it was possible to obtain an estimate not only of P_{CENSOR} , but also of the rms uncertainty in our answer. In Table IV, we list our results. These results are the basic objective of our numerical exercise. In the next section, we discuss the interpretation of these results.

Discussion of Results

The basic results are those presented in Table IV and we shall center our discussion about these results. The very large variation of P_{CENSOR} with D/r_0 is apparent from even a cursory examination. In Fig. 1, we have plotted P_{CENSOR} as a function of $(D/r_0)^2$. It is interesting to note how well the data is fit by an exponential dependence on aperture area. This is in

good agreement with an earlier conjecture by Hufnagel,⁸ though the coefficients of the fit are significantly different from those suggested by Hufnagel. We find that the data is well represented by the relationship

$$P_{CENSOR} = 5.6 \exp \left[-0.1557 \left(D/r_o \right)^2 \right] ,$$
 (55)

at least for values of $D/r_0 \ge 5$.

It is interesting to remark that if a CENSORING experiment were performed with $D/r_0 = 15$, as would be the nominal condition for a 1.5 m telescope (with ro nominally equal to 0.1 m), then the probability of getting a good picture in a single short exposure would be about 3.4 x 10-15. If independently distorted wavefront short exposures could be obtained at the rate of 100 per second, it would take more than 800 million hours "on an average" to get a good picture, i.e., one for which the average wavefront distortion over the aperture was less than one-radian. If the aperture diameter were reduced to 1 m, so that $D/r_0 = 10$, the probability would be about 1.1 × 10-6, and the expected waiting time to get a good picture would become about 2.5 hours (if we can get 100 independent wavefront distortion samples per second). With a 0.7 m diameter aperture, the waiting time shrinks to only 3.5 seconds. Clearly, in a CENSORING experiment, it is critical to know what ro is and to not make the aperture diameter much larger than about 7 ro, unless very long waiting times are acceptable.

We note that in certain cases, astronomical seeing with r_0 values in excess of 0.15 m have been reported. In such cases, it would be quite appropriate to attempt a CENSORING experiment with a 1 m diameter aperture, but it is critical that the aperture be properly stopped, and this requires current knowledge of r_0 , and appropriate planning in the implementation of the experiment.

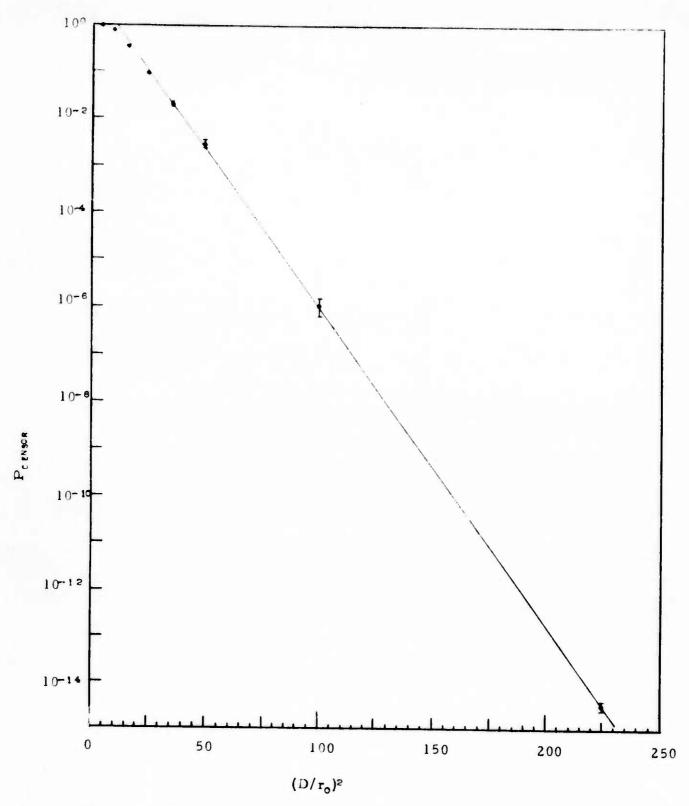


Figure 1. Probability of Obtaining a Good Short Exposure Image as a Function of Aperture Diameter. A good image is defined as one with less than one radian effective wavefront error (i.e., wavefront error excluding tilt) over the aperture. Aperture diameter D is measured in units of the wavefront distortion length, ro. Person is the probability.

Table III

Eigenvalue List

Each eigenvalue and its associated values of p and q are listed, for all values covered in Table II. When $q \neq 0$, the eigenvalue is considered to be listed twice, as indicated by the nature of the overall rank-order column, N, and by the Q-value column.

V		EIGEN VALUE	CUM. SUM.	P	つ
1 •	ے	.118765000	.037530000	1	22
3		.019733000	.055263000	i	0
4 •	5	.005215-00	.066694000	1	3, -3
5.	1	.005140000	.077054000	1	1, -1
선 •	4	.002153400	·091360900	1	44
10.	11	.001645400	.084652000	5	55
15		.001632500	·085284200	5	0
13.	14	· 001084400	. 04845 3800	3	55
15.	15	.000773650	.090001100	5	3, -3
17.	18	.000757140	.0915153HD	5	11
13.	20	.000617730	.092750840	1	66
21.	25	·000428150	.093607140	S	44
23.	24	.000382540	.094372720	1	77
25.	54	• U0 U3 H5 I 3 O	· 095136480	3	55
27		.000373790	.095510270	3	0
54.	29	.000565110	.096034490	5	55
30 •	31	.000251910	.096538310	1	AB
32.	33	.000224370	. 145987050	3	33
34.	37	.000215100	.097417250	3	11
38.	34	.000173460 .000172140	.097754970	1	9, -9
40.	41	.000143750	.098109050	2	66
42.	43	.0001334H0	.098396550 .098664310	3	44
44	45	.001129480	.098794190	4	25
45.	46	.00124550	.049043310	1	10 - 10
47.	48	.000119970	.099281050	ځ	1010
47.	50	.000097HR1	.099476812	3	55
51.	52	.000091463	·049660738	í	1111
53.	54	.000089079	.099638896	4	33
55.	56	.000085452	.100009900	2	AA
57.	54	.000094R0H	.100179416	4	11
59.	60	.000069736	.10031888R	3	66
61.	62	.000069546	·100458180	1	1212
63.	64	.00006339H	.100584976	2	3, -9
65.	66	.000062420	.100709816	4	44
67.	68	.000058784	.100827384	5	22
69		.000057999	.100885283	5	n
70.	71	.000053455	.100992993	1	1313
72.	73	.000051459	.101095911	3	77
74.	75	.00004R271	.101192453	5	1010
76.	77	.000045524	·101283501	4	55
78.	79	.000042494	.101368489	5	3, -3
80.	Al	.00004242A	•101453345	1	14,-14
82.	E 8	.000041356	.101536057	5	11
84.	85	.000039058	.101614173	3	AB
86.	A7	• 000037553	·101689279	5	1111
88.	89	.000034265	•101757R09	4	56
90.	93	.000033736	.1018254R1	1	1515
94	47	.000031690	.1018A9041	5	44
77		.000030739	.101919780	6	0

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P
  V
         EIGEN VALUE
                        CUY. 514.
                                           ()
 97. 95
          .000030374
                        .1019A045R
                                       3
                                           9. -9
 91. 9H
          .000029757
                        .102039472
                                          12 - 12
                        .102099440
 97.100
          .000029734
                                           2. -2
                        .102154520
101.102
          .000027540
                                          15,-16
                                       1
                        ·102207436
                                           7, -7
103-104
          .000026454
                        .102255958
          .000024261
                                       5
                                           5. -5
1030105
107-108
          .000024029
                        .102304016
                                       3
                                          10 -- 10
109.110
          .000023452
                                          13.-13
                        .102351920
111-112
          .000023465
                        ·102398850
                                       6
                                           1. -1
113.114
          295220000
                        .107444574
                                       6
                                           3. -3
                        .102489780
                                          17,-17
115.116
          .000022403
          .000020967
                        .102531514
                                           A. -B
117.118
                        .102570412
                                       5
          .000019549
                                          14,-14
117.120
151 - 155
          .000019348
                        .10260930R
                                          11.-11
                                           6. -6
123.124
                                       5
          .000019007
                        .102647322
                                          18,-18
                                       1
152+150
          .000018773
                        .1026A4A6A
127
          .000018529
                        .102703397
                                       7
                                           n
129.129
          .000017461
                        .102739119
                                       6
                                           4. -4
          .000016753
                        .102772625
                                           9. -9
130.171
                                       7
                                           5. -5
132.133
          . 000016643
                        ·102805911
                                       2
                                          15,-15
          .000016146
                        FOSRFRSOI.
134.175
                                          15.-15
          ·000015404
                        .102A69A11
                                       ٤
134.137
                                       1
                                          19.-19
          .000015725
134+139
                        .102901261
                                           7. -7
                                       5
          .000015192
140.141
                        .102931625
                                       7
                                           1. -1
142.143
          .000014H38
                        .102961301
                                           5, -5
144.145
          ·000014272
                        .102989705
                                       6
                                       4
                                          10,-10
145.147
          .00001365B
                        .103017021
          .000013482
                        ·103043985
                                       2
                                          16.-16
149.144
                                           3. -3
          .000013446
                        .103070977
150 - 151
          .000013307
                        .103097491
                                          20 -- 20
152.153
                                       1
154 - 155
                                          13.-13
          .000013071
                                       3
                        .103123633
                                       5
                                           R. - B
155 - 157
          .000012327
                        .10314A2A7
                                       R
154
          .0000172772
                        .103160519
                                           0
                                           6, -6
159.160
          .000011478
                        .1031A3475
                                       ħ
                                          17.-17
161.162
          .000011363
                        .103206201
                                       2
163.164
          .000011331
                        .103228863
                                          21 . - 21
165.166
          ·0000112H1
                        .103251425
                                          11.-11
          .000010931
                                       3
                                          14 .- 14
167.168
                        · 103273281
                                       7
                                           4. -4
          .000010492
164.170
                        .103295071
                                       A
                                           7. -2
171.172
          .000010318
                        .103315707
                                           3. -0
173.174
          .000010151
                        .103336009
1750176
          .000010135
                        .103356279
                                           1. -1
          .00000973H
                        .103375754
177.178
                                       1
                                          22 - - 22
          .000009462
                                       5
                                          14.-18
177.140
                        ·103395079
                                          12.-15
          .000009475
                        .103413934
141.185
                                           7. -7
143.144
          . 000009415
                        .103432763
                                       6
                                          15.-15
185.186
          .0000009230
                        .103451223
                                       3
                                       7
187.18H
          .0000009917
                                           5. -5
                        .103469056
          SEARUNOON.
                                       4
                                            n
189
                        .1034776HR
          .00000R4PH
190.191
                        .103494665
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FIGEN VALUE
  V
                         CUM. SIM.
                                      D
                                           1)
192.193
          .0000008464
                        .103511593
                                       5
                                          10.-10
194.195
          .0000000005
                                          23.-23
                        .103528403
                                       1
196.19/
          .00000A27A
                                          14.-19
                        .103544959
                                       2
194.199
          .000007759
                        .103560A76
                                       4
                                          17.-13
105.005
          .000007464
                        .103576604
                                       3
                                          16.-16
202.203
          .000007423
                        .103592250
                                       6
                                           H. -8
204.205
          .000007456
                                       9
                                           1. -1
                        .103607163
204.201
          .000007387
                                       7
                        .10362193A
                                           6. -6
405.805
          .00000731H
                                          24 . - 24
                        .103636574
                                       1
510
                                     10
          .000007152
                        .103643776
511.515
                        .103658016
                                          20.-20
          .000007145
213.214
          .000007134
                                       5
                        .1036727A4
                                          11.-11
215.216
          . 100007055
                                       A
                        .103686393
                                           4. -4
217.21H
          .000006405
                        .103700203
                                       9
                                           7. -2
219.220
          .000006792
                        .103713766
                                       4
                                          14 . - 14
221.222
          .000006751
                                          17.-17
                        .103727268
                                       3
223.224
          .000006576
                                           9. -9
                                       4
                        .103740470
225.226
          .000006435
                                     10
                        .103753200
                                           1 . -1
227.224
          . 000006 189
                        .103756067
                                          25 .- 25
                                      1
0529.530
          . 0000006205
                        .103778477
                                      3
                                          21 .- 21
231.232
                                           7. -7
          .000006192
                                       7
                        .103790R61
233
          .000006166
                        .103797027
                                     11
                                           0
234.235
                                          12.-12
          .000006071
                        .103A19169
234.231
          .000005911
                        .103R20990
                                      H
                                           5. -5
          .000005440
234.279
                        .103P32671
                                       3
                                          19.-18
240.241
          .000005425
                        .103R44321
                                          15,-15
242.243
                        .103855662
          .000005571
                                      9
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1043-1044	.000000381	.105057169	8	35,-35
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           .000000063
                         .105122277
                                       15
                                            32 . - 32
1365.1766
           .000000062
                                       14
                                            3A, -38
                         .105122401
1367-1368
           .000000061
                         .105122523
                                       20
                                             3. -3
1369.1370
           .000000058
                         .105122640
                                       19
                                             9. -17
1371.13/2
           .00000005H
                         .105122756
                                       18
                                            15 .- 15
1373.1374
          .000000058
                                       17
                         .105122972
                                            21 -- 21
1375.1376
          .000000058
                         .105122989
                                       16
                                            27 . - 27
1377.13/8 .000000058
                                       15
                         .105123105
                                            33,-33
1379.1380 .00000005A
                         .105123221
                                       14
                                            14.-39
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ELGEN VALUE
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                          .105123329
                                        14
                                            40 . - 40
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                          .105123437
                                        15
                                            34 . - 34
 1345.1304 .000000153
                          .105123542
                                       16
                                            28.-28
 136/-1384 .000000051
                          .105123645
                                       17
                                            22 - - 22
 1307-1390 .000000051
                          .105123747
                                       14
                                            41 . - 41
 1391.1392 .0000000000
                          .105123446
                                            15,-35
                                       15
 1393.1394 .000000049
                          .105123944
                                            14.-16
                                       18
 1395-1395 . 000000048
                          ·105124040
                                            29,-29
                                       16
 1307.1399 . 000000046
                          .105124132
                                       15
                                            36,-36
 1399.1400
            .0000000145
                          .105124224
                                            23,-23
                                       1/
 1401-1402 .000000044
                          .105124312
                                       19
                                            10.-10
 1403-1404 .000000044
                          .105124400
                                            30 .- 30
                                       16
 1405.1406 .000000043
                          .105124495
                                       15
                                            37 . - 37
 1407-1404 .0000000042
                          .105124569
                                       IA
                                            17.-17
 1409-1410 .000000041
                          .115124660
                                       11
                                            24 . - 24
 1411-1412 .000000140
                          .105124731
                                       16
                                            31 . - 31
 1413-1414 . 0000000000
                          .105124R11
                                       15
                                            34.-38
 1415-1415 .000000037
                          .105124AA5
                                       15
                                            34.-39
1417-1414 .000000003/
                          .105124959
                                       16
                                            32,-32
1417.1420 .000000037
                          .105125032
                                       17
                                           25 .- 25
1421-1422 .000000036
                          .105125104
                                       14
                                           19.-1H
1423-1424 .000000035
                         .105125173
                                       15
                                           411 -- 40
1425-1426 .000000034
                         .105125242
                                       19
                                           11.-11
1427.1429 .000000034
                         .105125310
                                       16
                                           33,-33
1427.1430 .0000000333
                         .105125376
                                       17
                                           26 . - 26
1431-1432 .000000033
                         .105125441
                                       15
                                           41 . - 41
1433.1434 .000000031
                         .105125504
                                       16
                                           34 . - 34
1435.1436 .000000031
                         .105125566
                                       18
                                           19,-19
143/-1434 .0000000330
                         .105125625
                                      17
                                           27 . - 27
1439.1447 .0000000029
                         ·105125683
                                      16
                                           35 . - 35
1441.1442 .000000028
                         .105125739
                                      20
                                            4. -4
1443.1444 .0000000077
                         .105125794
                                      19
                                           17.-12
1445.1446 .000000027
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                                      14
                                           20 . - 20
1447-1444 .000000007
                         .105125902
                                      16
                                           36,-36
1447.1450 .000000027
                         .105125956
                                      17
                                           24. -2H
1451 - 1452 . 000000025
                         .105126006
                                      16
                                           37,-37
1453-1454 .000000025
                         .105126055
                                      17
                                           29 . - 29
1455.1456 .000000024
                                           21,-21
                         .105126103
                                      18
1457.1454 .000000073
                         .105126149
                                      16
                                           JH . - 36
1459.1460 .000000022
                         .105126194
                                      17
                                           30 . - 30
1461-1402 .000000072
                         .10512623A
                                      19
                                           13.-13
1463.1464 .0000000025
                         .1151767A?
                                      16
                                           34.-39
1465.1466 .000000021
                        .105126324
                                      14
                                           55.-55
1467-1464 .000000021
                         .105126365
                                      17
                                           31 .- 31
1469.1470 .000000020
                        .105126406
                                      16
                                           40.-40
1471-1472 .000000019
                        .105126444
                                      16
                                           41 . - 41
1473-1474 .000000019
                        .105126482
                                      17
                                           32.-32
1475-14/4 .000000019
                        .105126519
                                      18
                                           23.-23
1477-14/H .00000001H
                        .105126555
                                      19
                                           14.-14
1479-1460 .000000017
                        .105126530
                                      17
                                           33.-33
```

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V
          FIGEN VALUE
                          CUM. SUM.
                                             3
1461-1442 .000000117
                         .105126623
                                            24 . - 24
                                       14
1483.1484 .000000 116
                                       17
                         .105126655
                                            14,-34
1485.1485 .000000015
                         .105126686
                                       20
                                             5, -5
1487.1488 .000000015
                                       19
                         .105126715
                                            15.-15
1487.1490 .000000015
                         .105126745
                                       18
                                            25 - 25
1491-1492 .000000015
                         . 105126775
                                       17
                                            35 . - 35
1493.1494 .000000014
                         .105126803
                                       17
                                            36 . - 36
1495.1496 .000000013
                         .105126829
                                       14
                                            26 - 26
1497 · 149H . 000000013
                         .105126855
                                       17
                                            37 . - 37
1499.1500 .000000013
                         .105126AA0
                                       14
                                            16 - 16
1501.1502 .000000012
                                       1 H
                         .105126905
                                            27.-27
1503-1504 .000000012
                         .105126924
                                       11
                                            34 . - 3H
1505-1505 .000000011
                         .105126951
                                       17
                                            39 - 39
                                       18
1507.1504 .000000011
                         .105126973
                                            28 - 28
                                       19
1509.1510 .000000011
                         .105126994
                                            17.-17
1511-1512 .000000010
                         .105127015
                                       11
                                            40,-40
1513-1514 .000000010
                         .105127035
                                       18
                                            24 . - 24
1515-1514 .0000000010
                         .105127055
                                       17
                                            41 . - 41
1517-1518 .0000000009
                                             6. -6
                         .195127073
                                       20
1513.1520 .000000009
                         .105127092
                                       19
                                            19.-18
1521.1522 .000000009
                         .105127110
                                       18
                                            30 . - 30
1523.1524 .0000000008
                                       1 H
                         .105127127
                                            31,-31
                                       19
1525-1526 .000000008
                         .105127143
                                            19.-19
1527.1528 .000000008
                         .105127158
                                       18
                                            32.-32
1523.1530 .0000000007
                         .105127172
                                       18
                                            33.-33
1531 • 1532 • 000000007
                                       19
                         .105127186
                                            50 -- 50
1533-1534 .0000000007
                                       18
                         .105127199
                                            34 . - 34
1535.1536 .000000006
                                       20
                         .105127212
                                             7. -7
1537.1538 .000000006
                         .105127224
                                       19
                                            21 . - 21
1539-1540 .000000000
                                       18
                         .105127236
                                            35 - 35
1541.1542 .0000000006
                         .105127247
                                       18
                                            36 - 36
                                       19
1543.1544 .000000005
                         .10512/258
                                            22 - - 22
1545 - 1545 - 000000005
                                       18
                         .10512726B
                                            37 . - 37
1547 - 1549 . 0000000005
                         .105127278
                                       18
                                            38 . - 38
1549.1550 .0000000005
                         .105127288
                                       19
                                            23 . - 23
1551 . 1552 . 0000000005
                         .105127297
                                       18
                                            39,-39
1553-1554 .0000000004
                         .10512/305
                                       20
                                             P. - R
1555.1556 .000000000
                                       18
                         .105127314
                                            40,-40
1557 - 1558 - 000000000
                         .105127322
                                       19
                                            24 -- 24
1559.1560 .0000000004
                         .105127330
                                       18
                                            41,-41
1561 - 1562
           .0000000004
                         ·105127338
                                       19
                                            25,-25
1563.1564 .0000000003
                         .105127345
                                       19
                                            26 . - 26
1565-1565 .0000000003
                                             9. -9
                         .105127351
                                       20
1567 - 1564
           .000000003
                                       19
                                            27 - - 27
                         .105127357
1564.15/0
           .000000003
                                       19
                                            24.-2B
                         .105127363
15/1-15/2 .0000000003
                                            29 - - 29
                                       19
                         .105127364
1573.1574
           2000000000
                                       20
                                            10.-10
                         .105127373
15/5.1576 .000000002
                         .105127377
                                       19
                                            30 . - 30
1577.1574 .000000002
                         .105127382
                                       19
                                            31 . - 31
1579.1580 .0000000002
                         ·105127386
                                       19
                                            32,-32
```

```
FIGEN VALUE
    V
                           CUM. SIM.
                                         P
                                             3
 1541.1582 .000000002
                          .1051273A9
                                        20
                                            11.-11
 1583-1584 .000000002
                          .10512/393
                                        19
                                            33.-33
 1545.1586
            .0000000002
                          .105127396
                                       19
                                            34,-34
 1547.1584
            .000000002
                          .105127399
                                        19
                                            35.-35
 1569-1590 .000000001
                          .105127402
                                       20
                                            12 .- 12
 1591 • 1592 • 000000001
                          .105127405
                                       19
                                            36,-36
 1593.1594 .0000000001
                          .10512740A
                                       19
                                            37,-37
 1595.1596 .000000001
                          .105127410
                                       19
                                            3A,-38
 1597-1598 .000000001
                          .105127413
                                       20
                                            13.-13
 1599.1600 .000000001
                          .105127415
                                       19
                                            39,-39
 1601.1602 .0000000001
                          .105127417
                                       19
                                            40 -- 40
 1603-1604 .0000000001
                          .105127419
                                       19
                                            41,-41
1605.1606 .000000001
                          .105127421
                                       20
                                            14 .- 14
1607-1604 .000000001
                          .105127423
                                       50
                                            15.-15
1607.1610 .00000001
                         .105127424
                                       20
                                            15,-16
1611-1612 .000000001
                         .105127425
                                       20
                                            17.-17
1613.1614 .000000000
                         .105127426
                                       20
                                            19.-18
1615.1615 .000000000
                         .105127427
                                       20
                                            19.-19
1617-1618 .000000000
                         .10512742A
                                       20
                                            50 .- 20
1619.1620
            .000000000
                         .105127429
                                       20
                                           21 . - 21
1621 - 1622
            .0000000000
                         .105127429
                                       20
                                           22 -- 52
1623.1624
            .0000000000
                         .105127430
                                       20
                                           23,-23
1625.1625 .000000000
                         .105127430
                                       20
                                           24 . - 24
1627.1628 .000000000
                         .105127430
                                       20
                                           25 .- 25
1629.1630 .000000000
                         .105127431
                                       20
                                           26 . - 26
1631 - 1632 . 0000000000
                                       20
                         .105127431
                                           27 . - 27
1633.1634 .000000000
                         .105127431
                                       20
                                           29.-2A
1635.1636
           .0000000000
                         .105127432
                                       20
                                           29 . - 29
1637.163H
           .0000000000
                         .105127432
                                       20
                                           30 -- 30
1639.1640
           .0000000000
                         .105127432
                                      20
                                           31 . - 31
1641.1642 .0000000000
                         .105127432
                                      20
                                           32 . - 32
1543.1644 .000000000
                         .105127433
                                      20
                                           37,-33
1545.1645 .000000000
                         .105127433
                                      20
                                           34,-34
1547.1544 .000000000
                         .105127433
                                      20
                                           35.-35
1544.1650 .000000000
                         .105127433
                                      20
                                           34 . - 36
1651.1652 .0000000000
                         .105127433
                                      20
                                           37 . - 37
1653.1654 .000000000
                         .105127433
                                      20
                                           34,-38
1653.1656 .000000000
                         .105127434
                                      20
                                           39 . - 39
1657-1654 .000000000
                        .105127434
                                      20
                                           40 . - 40
1557.1660 .0000000000
                        .105127434
                                      20
                                           41 . - 41
```

Table IV

CENSORING System Probabilities

 P_{CENNOR} is the probability that an aperture of diameter D/r_0 will, during a single short exposure, receive a wavefront whose rms distortion over the aperture (with tilt not considered a form of distortion) will be less than one radian.

D/r _o	${ m P}_{{ m CENSOR}}$		
2	0.986 ± 0.006		
3	0.765 ± 0.005		
4	0.334 ± 0.014		
5	$(9.38 \pm 0.33) \times 10^{-2}$		
6	$(1.915 \pm 0.084) \times 10^{-2}$		
7	$(2.87 \pm 0.57) \times 10^{-3}$		
10	$(1.07 \pm 0.48) \times 10^{-6}$		
15	$(3.40 \pm 0.59) \times 10^{-15}$		

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